

The Virtual Observatory: A New Revolutionary Way of Scientific Research

Astroinformatics, e-Science

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Astroinformatics and VO Summer School
Faculty of Mathematics, University Belgrade, 30-th June 2010, Serbia

Outline of the Talk

- VO – the hidden revolution in astronomy
- Data Avalanche in astronomy
- History of VO
- Basic principles of technology
- VO Tools
- Theory in VO
- VO Science
- VO and Society

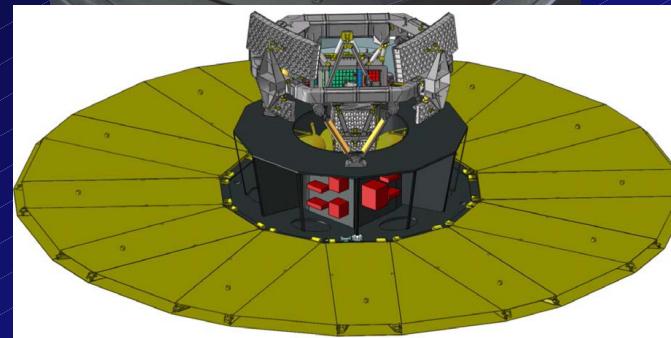
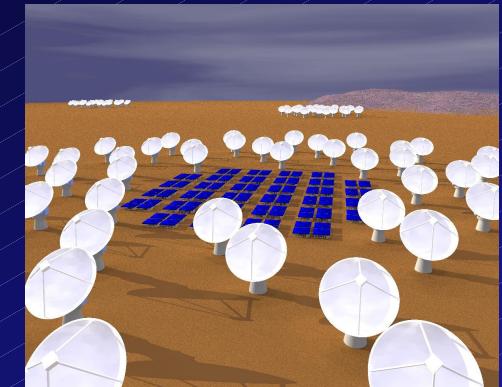
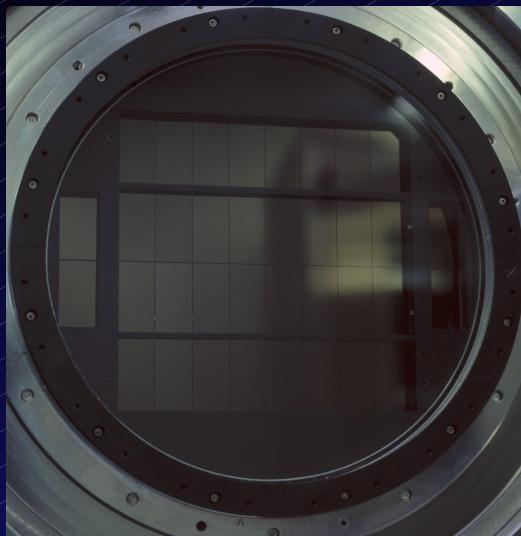
VO - The Hidden Revolution

- VO is the radical change of the paradigm of the work of the scientists – effectiveness !!!
- Everyday question (what, where, format, units)
- Everyone is using it – but not stated (> 5 years)
 - CDS (Simbad, Aladin, Vizier), NASA, ESA archives
 - All looks like „ONLY“ another WEBS, client apps
- Scientists are conservative (don't like change)
 - The fear of buzzword VO (multispec, large scale)
 - Computer literacy – obligatory (part of job)

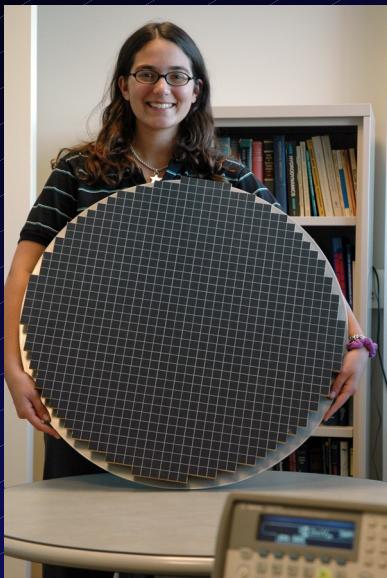
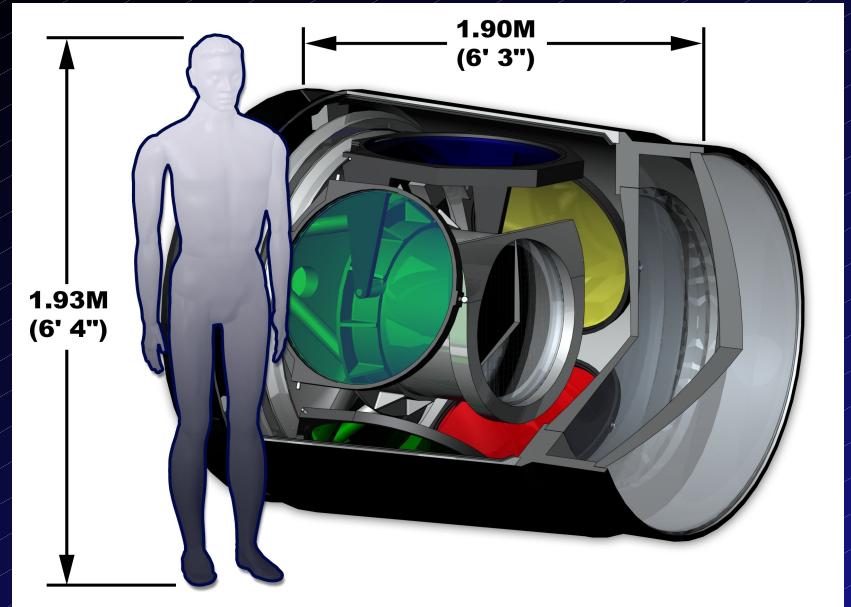
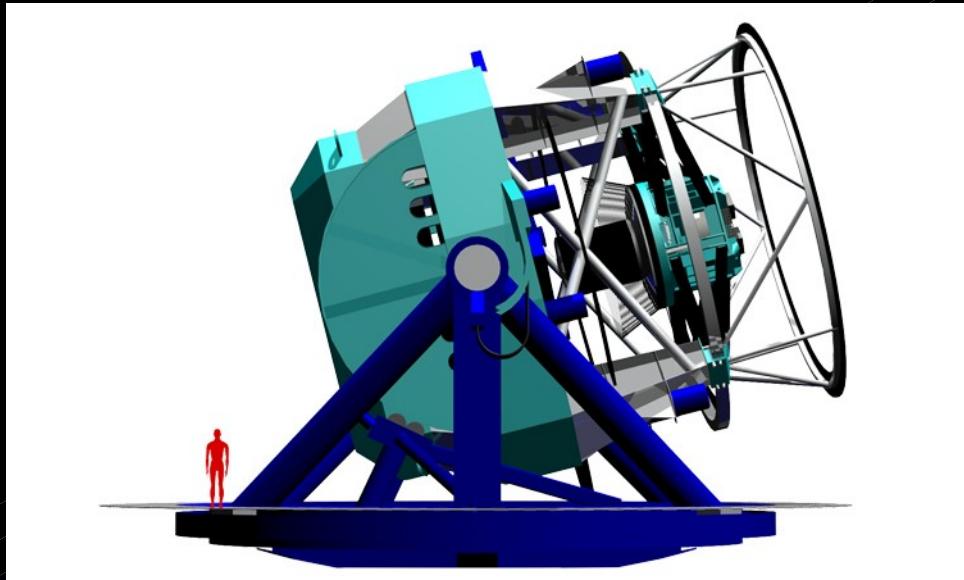
Analogy between VO and WWW

- Linking HYPERTEXT/DATA among servers
- Synergy effect of GLOBAL NET (Gopher,WAIS)
- Powerfull SEARCH (VERONICA – GOOGLE)
- DISTRIBUTED but CENTRAL Steering Organisation (W3C/IVOA)
- Recommendations = „Obligatory“ Standards
- Astronomers in forefront of development
- Scepticism (usefullness for my field ???)
- Steep Growth – average user can use it without knowledge of principles (effectivity, habits)

Data Avalanche



LSST (8.4m)

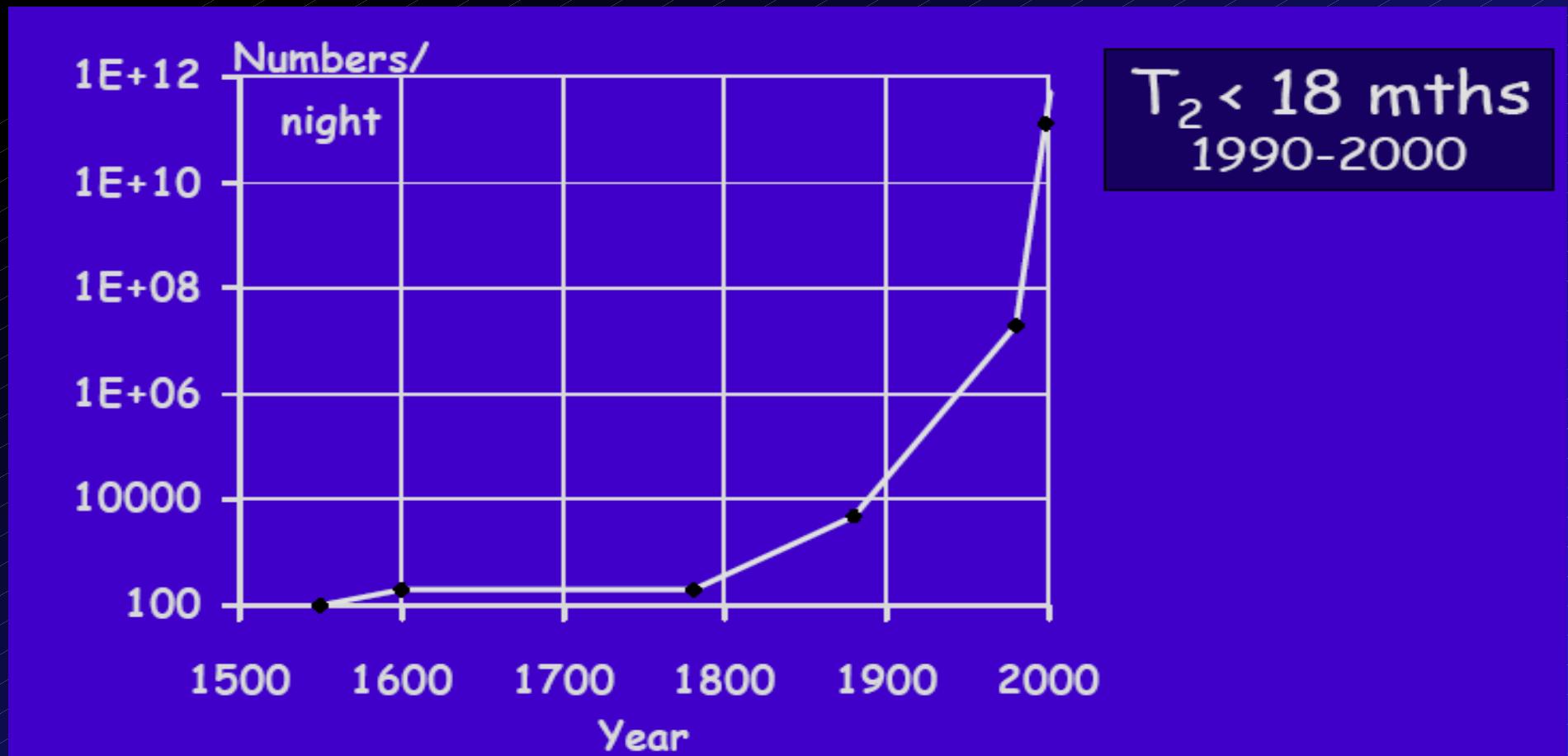


200 CCD 4kx4k,
32 channels (6400)
3.2 Gpix every 20 sec
64cm diameter
3.5 deg FOV
30 TB/night
2 TFLOPS
detection of changes
within 60sec

Data Avalanche

Moore law for chips –doubling 1.5 year

Data in astronomy – doubling < 1 yr ! (1000/10 yr)



History of VO

Success of IUE/HST archives

idea of the VO - end 2000

Federation of archives (MAST, NED)

unified IF, data format for transport

Huge data – distributed processing

GRID - started in HEP (accelerator science)

Multispectral research : radio---gamma

Virtual Universe (UK), AstroVirTel (ESO)

Data for SDSS, SIMBAD, NED – key research

Virtual Observatory : Key Definitions

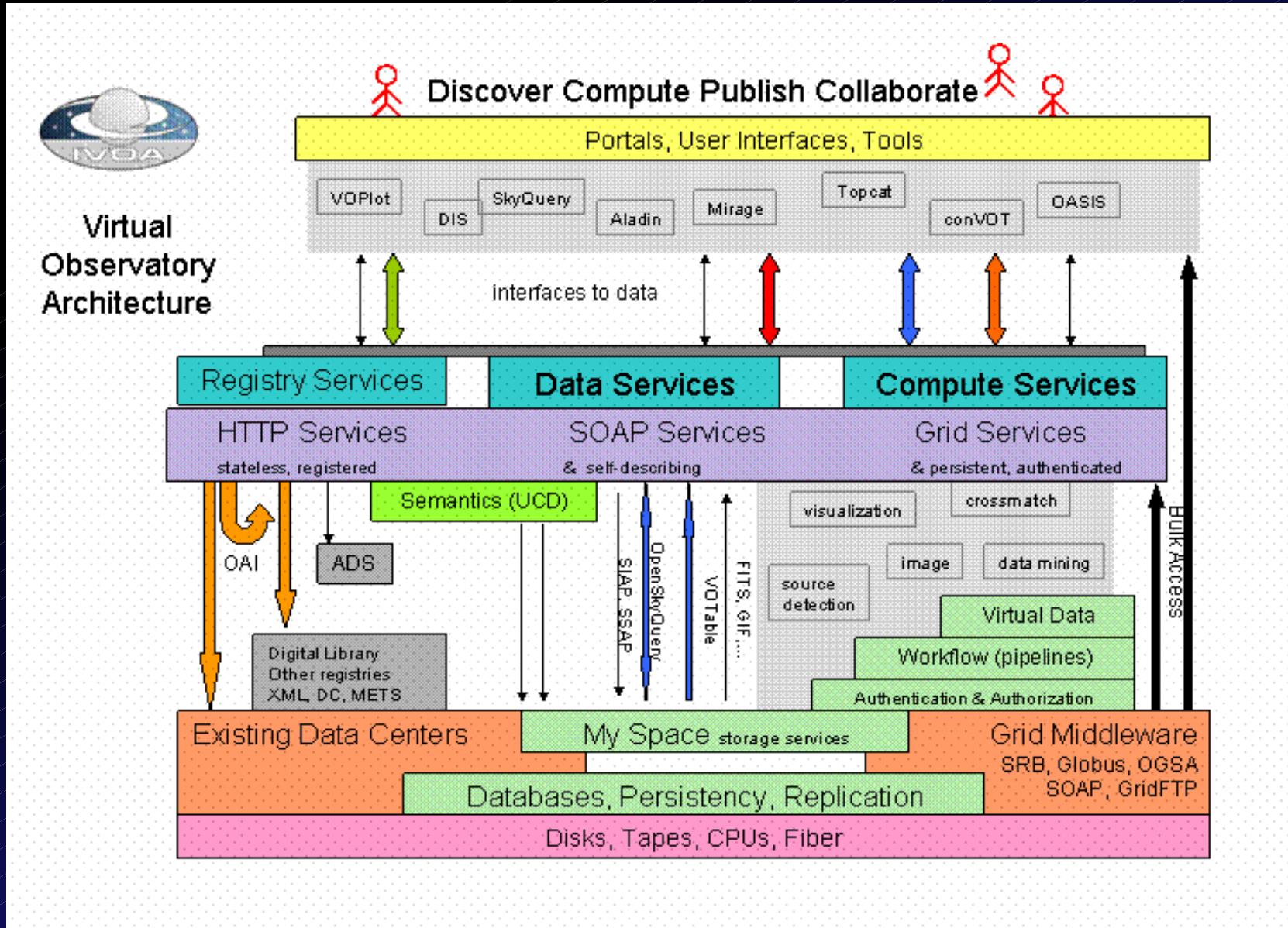
- “*The Virtual Observatory will be a system that allows astronomers to interrogate multiple data centers in a seamless and transparent way, which provides new powerful analysis and visualization tools within that system, and which gives data centers a standard framework for publishing and delivering services using their data*”.
- Standardization of data and metadata, and of data exchange methods.
- Registry, listing available services and what can be done with them.

R.J.Hanisch, P.J.Quinn, in “IVOA – Guidelines for participation”

VO Paradigma

- METADATA (name of column), ontologies (name)
- Unique format (VOTable – e.g Vizier)
- Transparent search, download, conversion
- Query for data – processing done on servers
- Federation of astronomical archives (protocols)
- Unified presentation – automatic units conversion
(A,MeV,MHz->nm), $\text{Wm}^{-2}\text{s}^{-1}$ → Jy)
- Background computing on GRIDS
- Multiwavelength approach (SED)

Architecture of VO



Technology of VO

Unified data format– VOTable, UCD (Vizier)

Transparent transport (SOAP , REST_(youtube))

Web services (WS) e-commerce, B2B, J2EE, .Net

VOregistry (DNS like) Google for data+WS
protocols (CGI)

ConeSearch (searching in circle on sky)

SIAP (Simple Image Access Protocol)

SSAP(Simple Spectral Access Protocol)

SLAP(Simple Line Access Protocol)

TAP (Table Access Protocol)

VOEVENT (transients, robotic telescopes,Sun)

VO Registry - web

Soubor Úpravy Zobrazit Historie Záložky Nástroje Nápověda

http://nvo.stsci.edu/vor10/getRecord.aspx?id=ivo://asu.cas.cz/stel/heros/cutout

Nejnavštěvovan... Getting Started Latest Headlines ELIAV, a.s. - Firmy.cz

VOSpec_Oct2007.png (PNG obrázek) HEROS OND CUTOUT: Resource Record Summary

NVO Registry

Hosted By Space Telescope Science Institute

Resource Record Summary

Catalog Service:

cutout server of HEROS archive of Ondrejov observations

Short name: HEROS OND CUTOUT
IVOA Identifier: ivo://asu.cas.cz/stel/heros/cutout
Publisher: Stellar Department of Astronomical Institute of the Academy of Sciences of the Czech Republic [+] [Pub. ID]
More Info: <http://stelweb.asu.cas.cz/vo-archives/heros>

Status: active
Registered: 2008 Oct 24 21:13:08Z
[Get XML](#)

Description

This is the cutout server for the content of HEROS archive. Using the parametr BAND=I1/I2 prepares on the fly spectra extended only over given range of wavelengths. This archive contains about 6000 high resolution (R=2000) echelle spectra obtained by the HEROS spectrograph installed at the 2m telescope of the Ondrejov observatory since August 2000 to March 2003. The spectra exposed simultaneously in two channels (red 580-840nm, blue 370-560nm) were reduced by MIDAS HEROS pipeline including the merging of echelle orders, heliocentric correction and rebinning to steps of 0.1A. The intensities are in instrumental flux (not normalized).

More About this Resource

[+] **About the Resource Providers**
This section describes who is responsible for this resource.

[+] **Status of This Resource**
This section provides some status information: the resource version, availability, and relevant dates.

[+] **What This Resource is About**
This section describes what the resource is, what it contains, and how it might be relevant.

Available Service Interfaces

[+] **Simple Spectral Access**
This is a standard IVOA service for searches for spectra from this resource that were observed within a specified region of the sky.

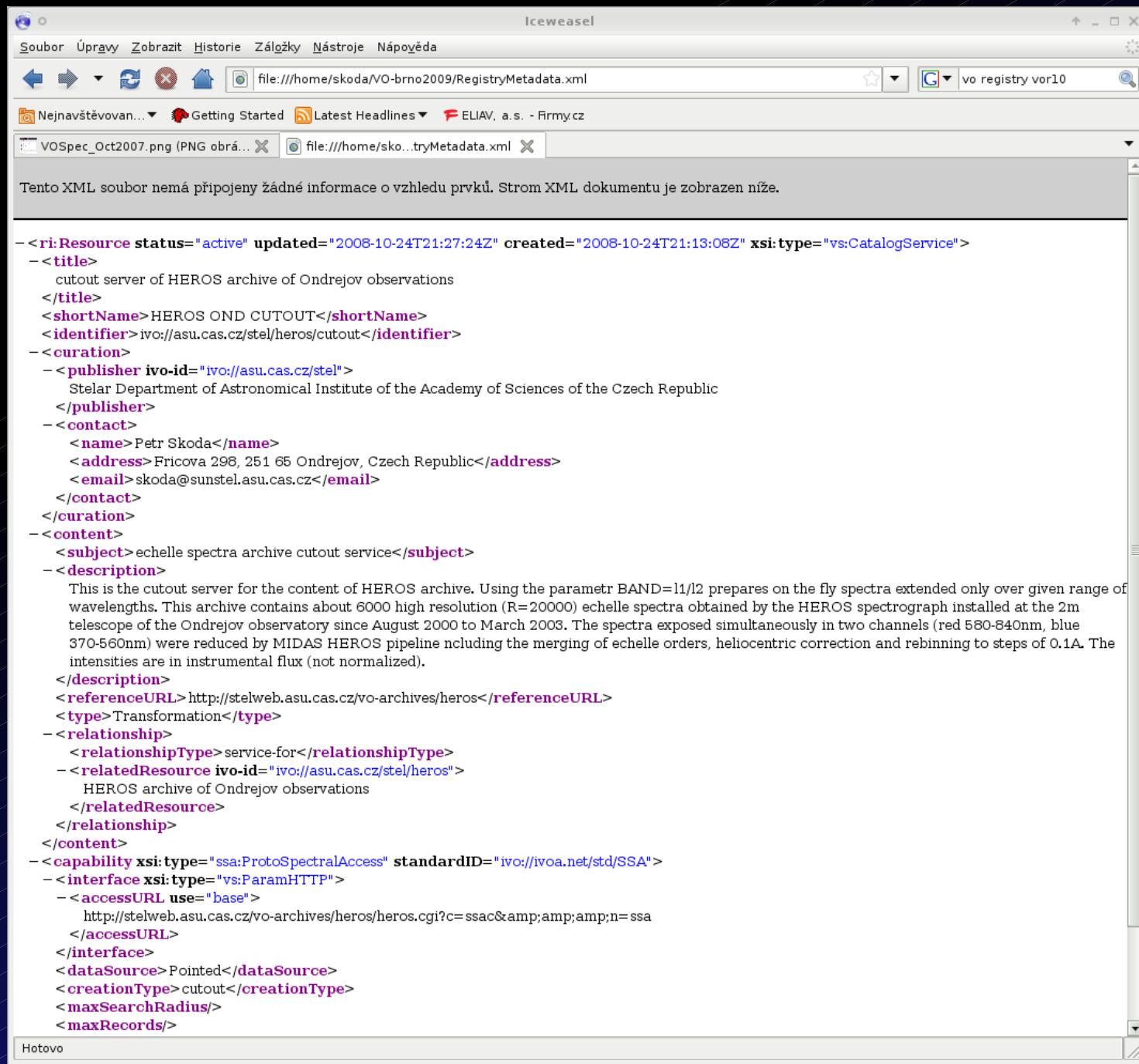
Developed with the support of the National Science Foundation under Cooperative Agreement AST0122449 with the Johns Hopkins University. The NVO is a member of the International Virtual Observatory Alliance.

This NVO Application is hosted by the Space Telescope Science Institute

Member

Hotovo

VO Registry - XML



The screenshot shows the Iceweasel web browser displaying the XML content of a VO Registry resource. The title bar reads "Iceweasel". The address bar shows the URL "file:///home/skoda/VO-brno2009/RegistryMetadata.xml". The main content area displays the XML code for the "HEROS OND CUTOUT" service.

```
<ri:Resource status="active" updated="2008-10-24T21:27:24Z" created="2008-10-24T21:13:08Z" xsi:type="vs:CatalogService">
  <title>
    cutout server of HEROS archive of Ondrejov observations
  </title>
  <shortName>HEROS OND CUTOUT</shortName>
  <identifier>ivo://asu.cas.cz/stel/heros/cutout</identifier>
  <curation>
    <publisher ivo-id="ivo://asu.cas.cz/stel">
      Stelar Department of Astronomical Institute of the Academy of Sciences of the Czech Republic
    </publisher>
    <contact>
      <name>Petr Skoda</name>
      <address>Fricova 298, 251 65 Ondrejov, Czech Republic</address>
      <email>skoda@sunstel.asu.cas.cz</email>
    </contact>
  </curation>
  <content>
    <subject>echelle spectra archive cutout service</subject>
    <description>
      This is the cutout server for the content of HEROS archive. Using the parametr BAND=l1/l2 prepares on the fly spectra extended only over given range of wavelengths. This archive contains about 6000 high resolution (R=20000) echelle spectra obtained by the HEROS spectrograph installed at the 2m telescope of the Ondrejov observatory since August 2000 to March 2003. The spectra exposed simultaneously in two channels (red 580-840nm, blue 370-560nm) were reduced by MIDAS HEROS pipeline including the merging of echelle orders, heliocentric correction and rebinning to steps of 0.1A. The intensities are in instrumental flux (not normalized).
    </description>
    <referenceURL>http://stelweb.asu.cas.cz/vo-archives/heros</referenceURL>
    <type>Transformation</type>
    <relationship>
      <relationshipType>service-for</relationshipType>
      <relatedResource ivo-id="ivo://asu.cas.cz/stel/heros">
        HEROS archive of Ondrejov observations
      </relatedResource>
    </relationship>
  </content>
  <capability xsi:type="ssa:ProtoSpectralAccess" standardID="ivo://ivoa.net/std/SSA">
    <interface xsi:type="vs:ParamHTTP">
      <accessURL use="base">
        http://stelweb.asu.cas.cz/vo-archives/heros/heros.cgi?c=ssac&amp;amp;n=ssa
      </accessURL>
    </interface>
    <dataSource>Pointed</dataSource>
    <creationType>cutout</creationType>
    <maxSearchRadius/>
    <maxRecords/>
  </capability>

```

Technology of VO

ADQL (Astronomical Data Query Language)

XMATCH, REGION (2 catalogues - shifted)

Application interoperability – PLASTIC, SAMP

Allows develop applications as bricks

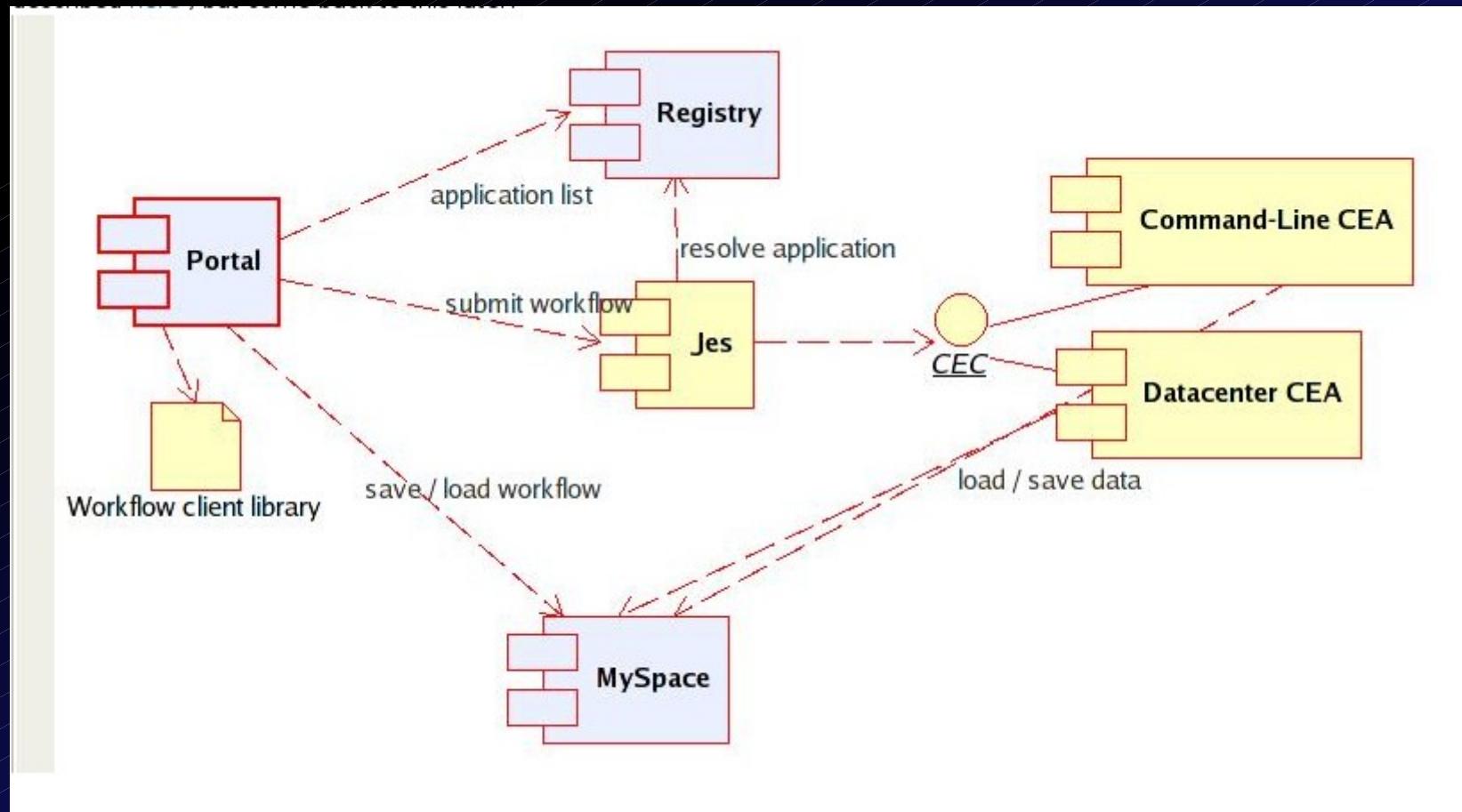
sending VOTABLES (catalogue-spectra-images)

Commercial interest (GoogleSky, MS WWT)

Planetariums, Outreach (Stellarium)

Workflows - Astrogrid

Running remote services – e.g. Sextractor, CASJobs, AstroNeural MLP....



Building a query

AstroGrid Query Builder

The screenshot illustrates the AstroGrid Query Builder interface for SDSS Data Release 3 (AstroGrid DSA). The interface is divided into several panels:

- Top Bar:** SDSS Data Release 3 (AstroGrid DSA) with icons for Chooser, Query, Parameter, XML, Info, and Security.
- Left Panel:** Chooser (selected), Query, Parameter, XML, Info, Security. Sub-tabs include Tree, Adql/s, and Adql/x. The Tree tab shows a hierarchical query structure:
 - Select
 - Items: a.ra, a.dec, a.u, a.g, a.r, a.i, a.z, a.err_u, a.err_g, a.err_r, a.err_i, a.err_z, a.type
 - From: PhotoObj as a
 - Where
 - And
 - a.ra Between 242.0 And 243.6
 - a.dec Between 54.1 And 55.1
- Table Columns:** A panel showing the columns of the PhotoObjAll table, which maps primary and secondary objects to a view. It includes columns: Name, UCD, Units, Type, and Description.
- Table Description:** A panel providing a detailed description of the PhotoObjAll table.
- Selected table:** A panel listing various tables in the database, with PhotoObj selected.
- Query being built:** A panel at the bottom left showing the current state of the query build.
- Diagnostics:** A panel at the bottom left for monitoring the query's performance.
- Dialog to insert selected columns into selected section of the query:** A yellow callout pointing to the 'Column References' section of the query tree, indicating where columns are being inserted.
- List of tables in the database:** A yellow callout pointing to the list of tables on the right side of the interface.

**VOTable :**

```
<?xml version="1.0"?>
<!DOCTYPE VOTABLE SYSTEM "http://us-vo.org/xml/VOTable.dtd">
<VOTABLE >
  <DESCRIPTION>
    VizieR Astronomical Server: urania.iucaa.ernet.in          2002-10-04T05:20:16
    Explanations and Statistics of UCDs:                      See LINK below
    In case of problem, please report to: question@simbad.u-strasbg.fr
  </DESCRIPTION>
  <DEFINITIONS>
    <COOSYS ID="J2000" equinox="J2000" system="EQ_FK5"/>
  </DEFINITIONS>
  <INFO ID="Ref" name="-ref" value="VOTx11451"/>
  <RESOURCE name="V105" ID="yCat_5105" >
    <DESCRIPTION>SKY2000 Catalog, Version 3 (Myers+ 2000)
    </DESCRIPTION>
    <TABLE ID="V_105_sky2v3r1" name="V105/sky2v3r1" >
      <DESCRIPTION>The Sky2000 Version 2 Catalogue
      </DESCRIPTION>
      <FIELD datatype="int" width="6" name="HD" ucd="ID_ALTERNATIVE" >
        <DESCRIPTION>Henry Draper &lt;math>1/35</math> number
        </DESCRIPTION>
      </FIELD>
      <FIELD unit="h:m:s" datatype="char" ref="J2000" name="RAJ2000" ucd="POS_EQ_RA_MAIN" arraysize="13" >
        <DESCRIPTION>Right ascension (J2000) hours
        </DESCRIPTION>
      </FIELD>
      <FIELD unit="d:m:s" datatype="char" ref="J2000" name="DEJ2000" ucd="POS_EQ_DEC_MAIN" arraysize="13" >
        <DESCRIPTION>Declination degrees (J2000)
        </DESCRIPTION>
      </FIELD>
```

 **Display Data Of Selected Points****Close****Save As File**

IVOA

Astro
Grid



EUROVO



Simple Spectra Access Protocol Spectral Data Model

Simple Spectral Access Protocol V1.04



*International
Virtual
Observatory
Alliance*

Simple Spectral Access Protocol
Version 1.04
IVOA Recommendation Feb 01, 2008

This version:
<http://www.ivoa.net/Documents/REC/DAL/SSA-20080201.html>

Latest version:
<http://www.ivoa.net/Documents/latest/SSA.html>

Previous version(s):
Version 1.03, December 2007
Version 1.02, September 2007
Version 1.01, June 2007
Version 1.00, May 2007
Version 0.97, November 2006
Version 0.96, September 2006
Version 0.95 May 2006
Version 0.91 October 2005
Version 0.90 May 2005

Editors:
D.Tody, M. Dolensky

Authors:
D.Tody, M. Dolensky, J. McDowell, F. Bonnarel, T.Budavari, I.Busko, A. Micol, P.Osuna, J.Salgado, P.Skoda, R.Thompson, F.Valdes, and the data access layer working group.



*International
Virtual
Observatory
Alliance*

IVOA Spectral Data Model
Version 1.03
IVOA Recommendation 2007-10-29

This version (Recommendation Rev 1)
<http://www.ivoa.net/Documents/REC/DM/SpectrumDM-20071029.pdf>

Latest version:
<http://www.ivoa.net/Documents/latest/SpectrumDM.html>

Previous versions:
<http://www.ivoa.net/Documents/PR/DM/SpectrumDM-20070913.html>

Editors:
Jonathan McDowell, Doug Tody

Contributors:
Jonathan McDowell, Doug Tody, Tamas Budavari, Markus Dolensky, Inga Kamp, Kelly McCusker, Pavlos Protopapas, Arnold Rots, Randy Thompson, Frank Valdes, Petr Skoda, and the IVOA Data Access Layer and Data Model Working Groups.

SSAP Parameters

4.1.1 Mandatory Query Parameters

The following parameters **must** be implemented by a compliant service:

Parameter	Sample value	Physical unit	Datatype
POS	52,-27.8	degrees; defaults to ICRS	string
SIZE	0.05	degrees	double
BAND	2.7E-7/0.13	meters	string
TIME	1998-05-21/1999	ISO 8601 UTC	string
FORMAT	votable	-	string

4.1.2 Recommended and Optional Query Parameters

Parameter	Sample value	Unit	Req	Datatype
APERTURE	0.00028 (=1")	degrees	OPT	double
SPECRP	2000	$\lambda/d\lambda$	REC	double
SPATRES	0.05	degrees	REC	double
TIMERES	31536000 (=1yr)	seconds	OPT	double
SNR	5.0	dimensionless	OPT	double
REDSHIFT	1.3/3.0	dimensionless	OPT	string
VARAMPL	0.77	dimensionless	OPT	string
TARGETNAME	mars		OPT	string
TARGETCLASS	star		OPT	string
FLUXCALIB	relative		OPT	string
WAVECALIB	absolute		OPT	string
PUBDID	ADS/col#R5983		REC	string
CREATORDID	ivo://auth/col#R1234		REC	string
COLLECTION	SDSS-DR5		REC	string
TOP	20	dimensionless	REC	int
MAXREC	5000		REC	string
MTIME	2005-01-01/2006-01-01	ISO 8601	REC	string
COMPRESS	true		REC	boolean
RUNID			REC	string

VO-enabled tools

Aladin

VOPlot

TOPCAT

VOSpec

SpecView

SPLAT

ViSiVO (HPC simulations, cosmology)

VOSED

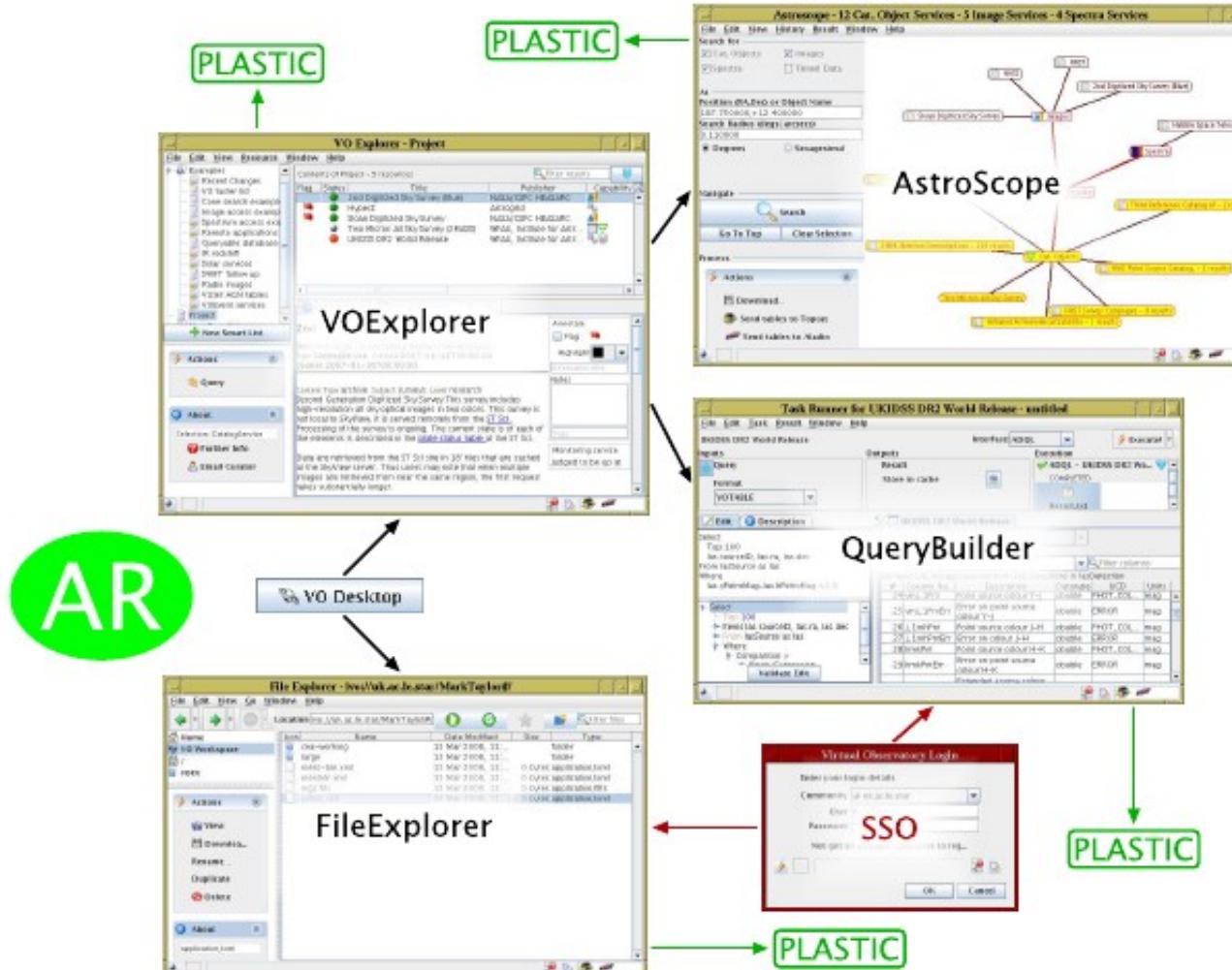
BASTI

SExtractor – WESIX (Web Enabled Source Identification with
Cross Matching)

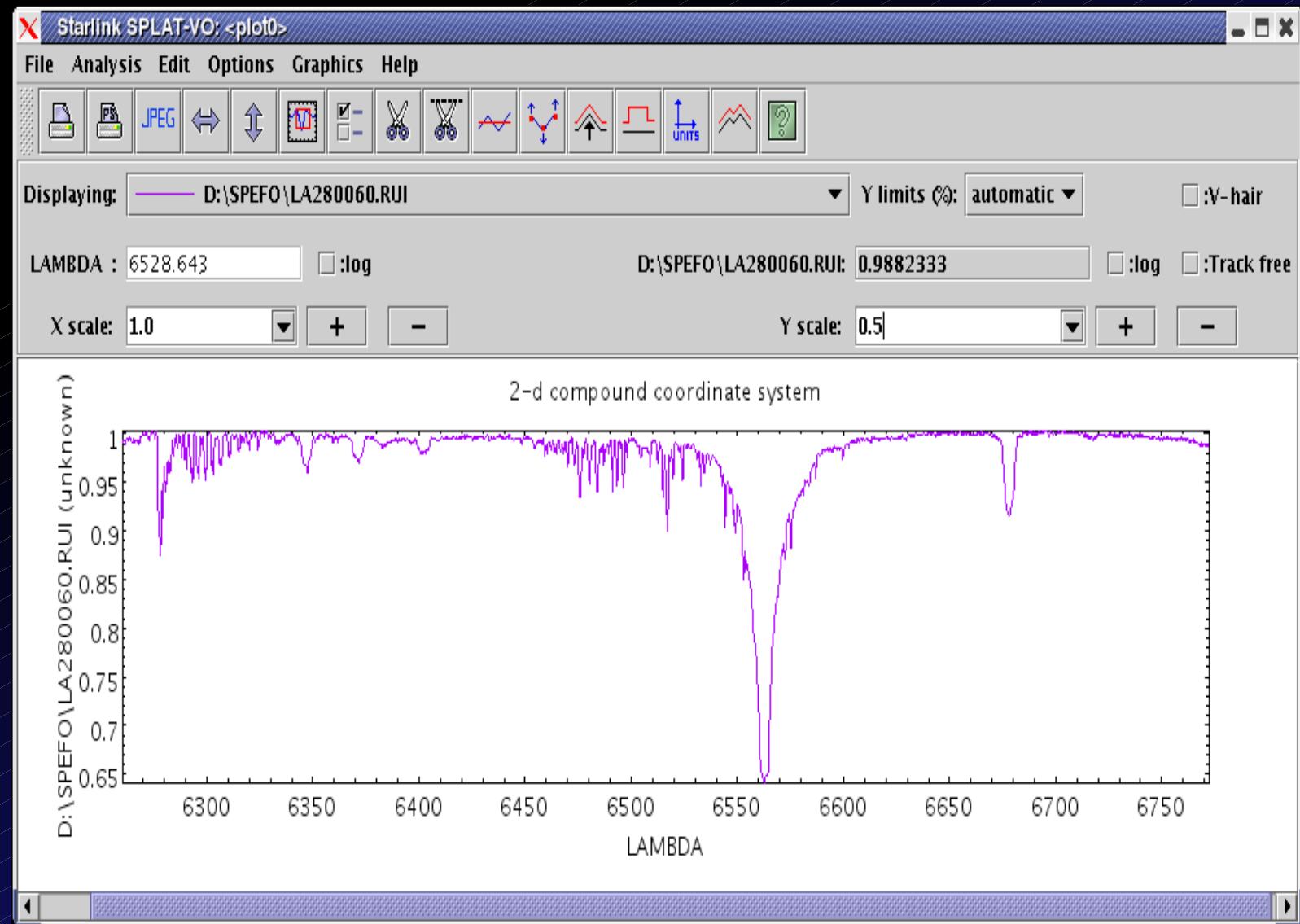
Period04 (since 18.9.08) - PLASTIC

AstroGRID VODesktop

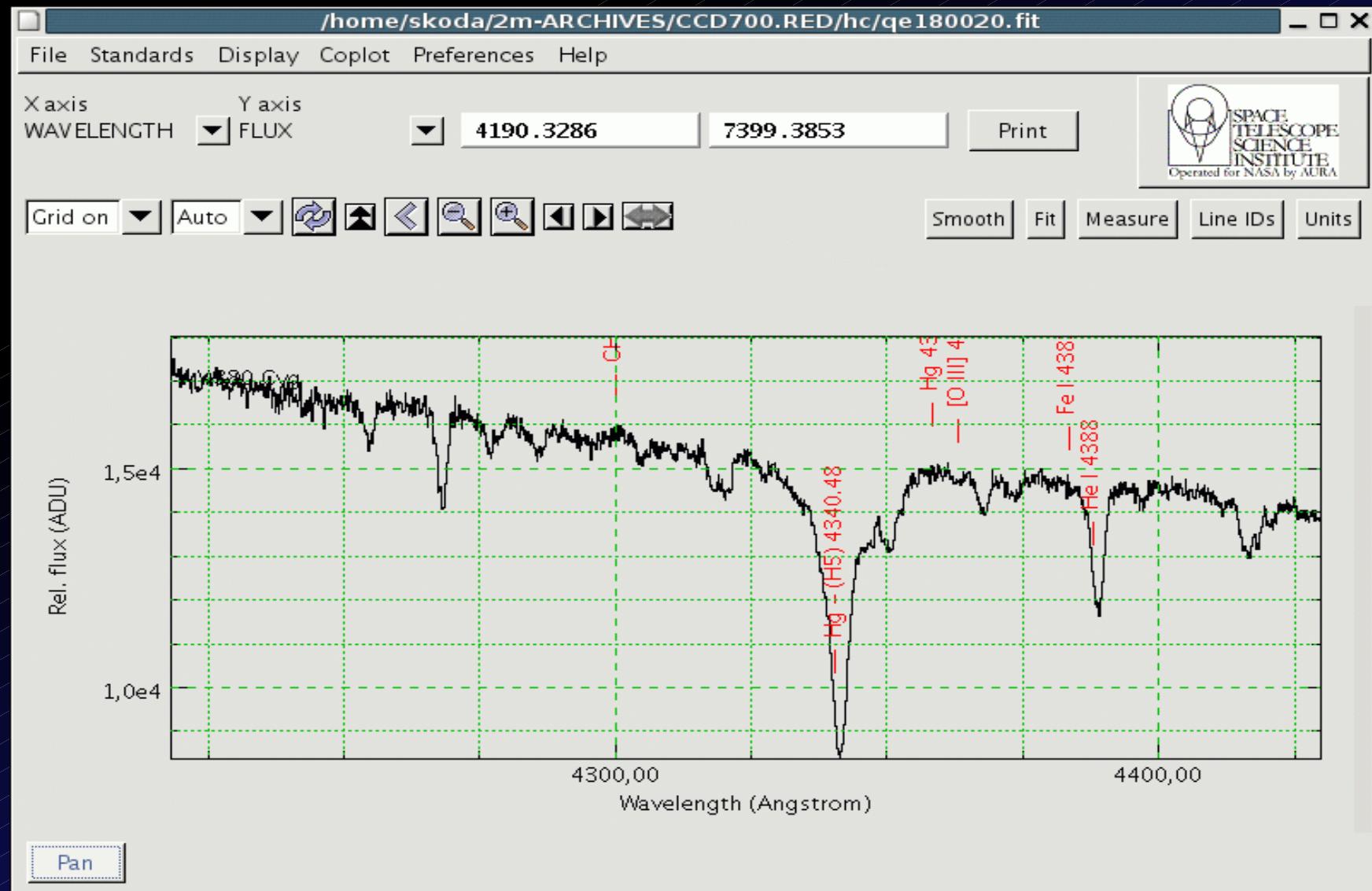
VODesktop Overview



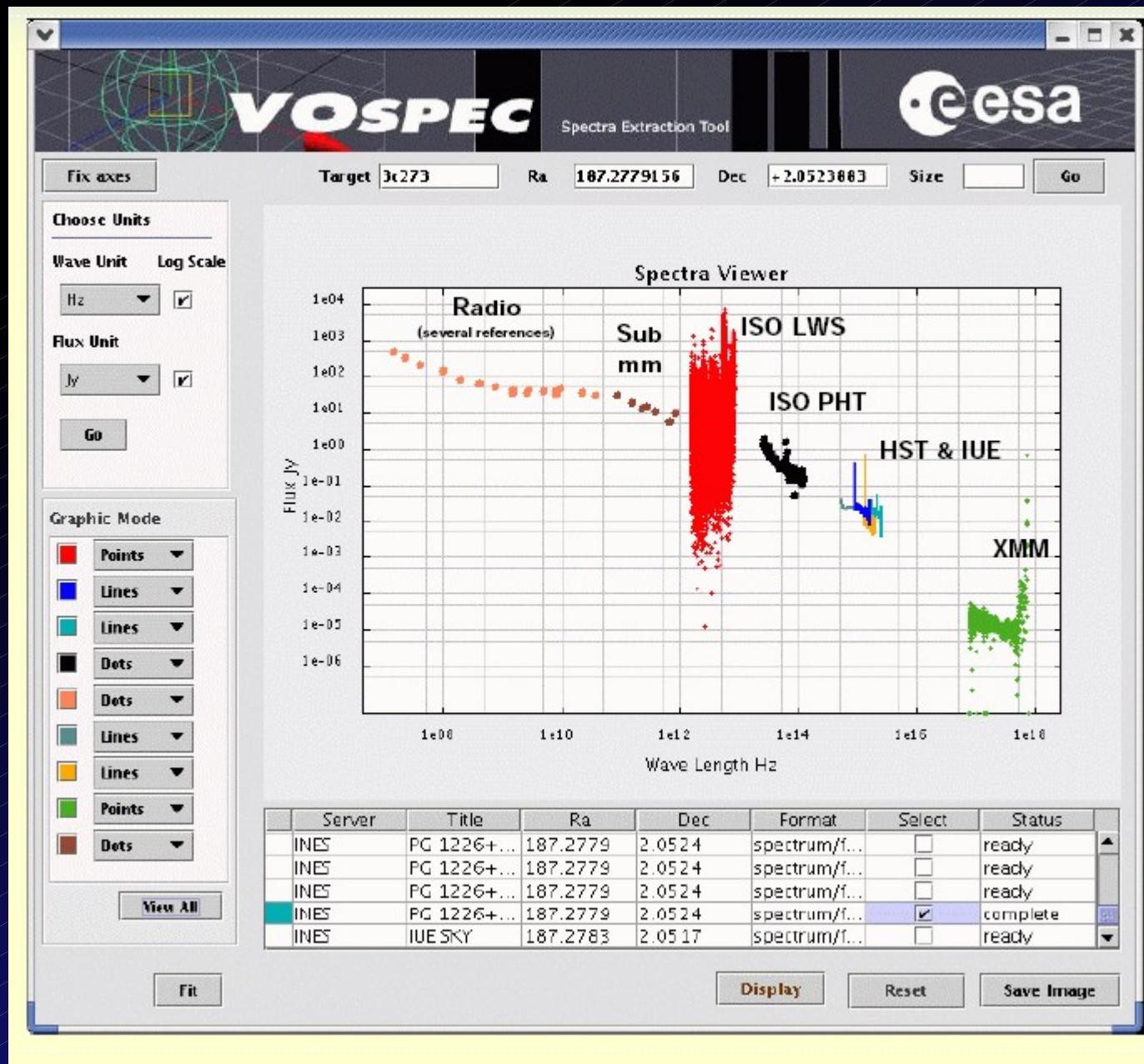
SPLAT-VO (Starlink, JAC)



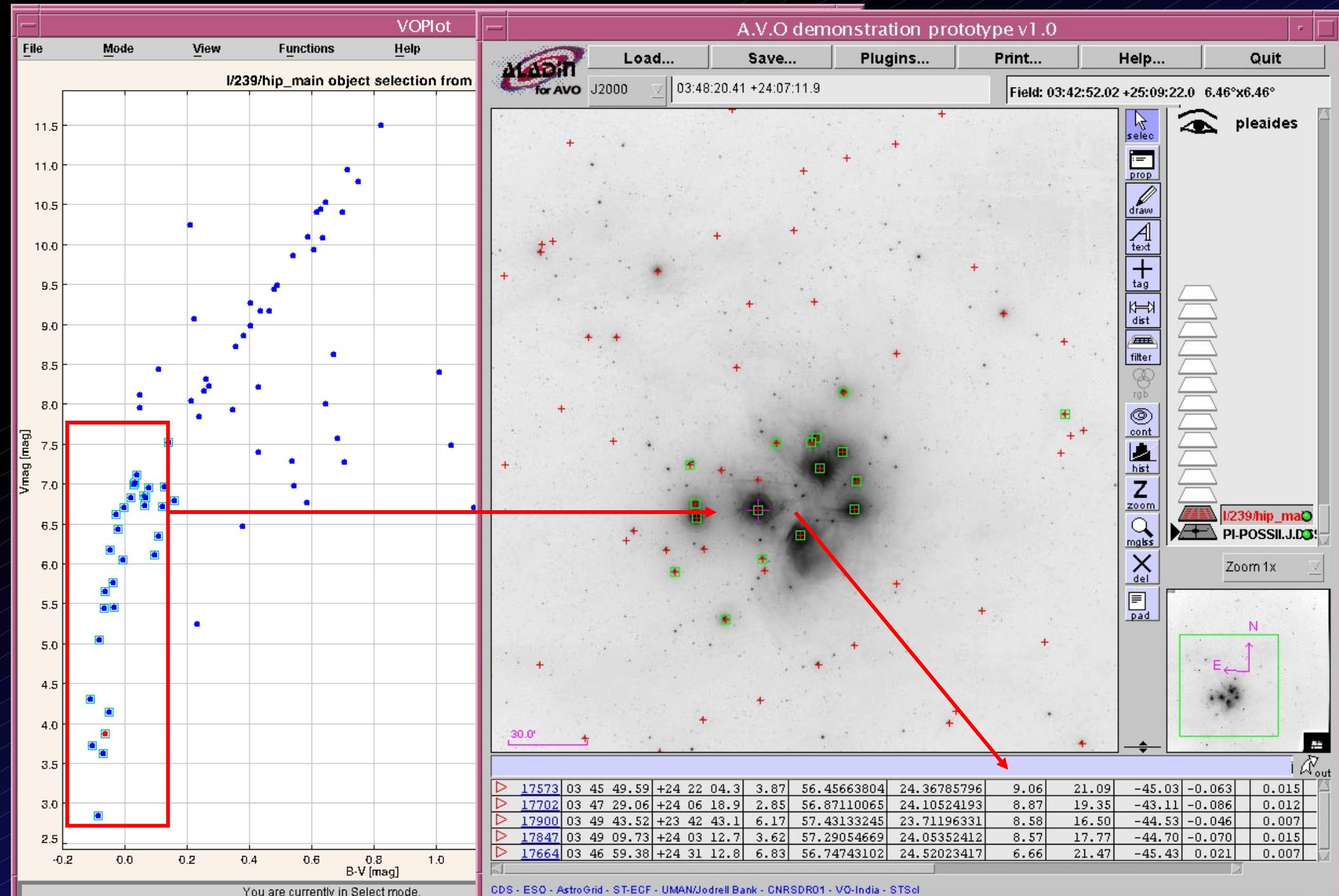
SpecView (STScI)



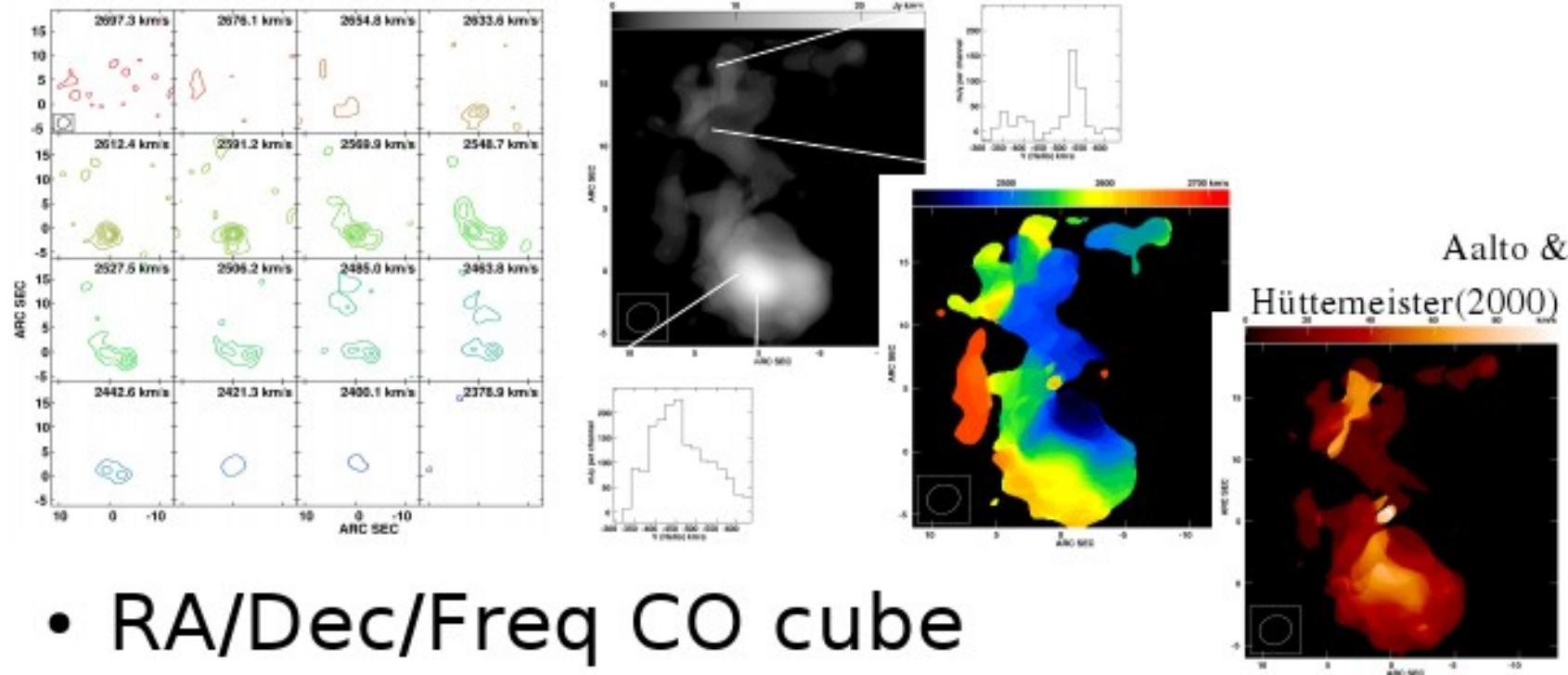
VOspec (ESAC)



Colour-magnitude diagram



ALMA/IRAM use case



- RA/Dec/Freq CO cube
 - Convert to velocity (LSR, radio convention)
 - Cutouts, simple squashes - VO tools?
 - Smoothed spectra, moments with noise cut-off
 - Specialised server-side pipeline controlled via UWS

Other VOs

Virtual Solar Observatory

Virtual Solar-Terrestrial Observatory

Virtual Magnetospheric Observatory

Virtual Space Physics Observatory

Virtual Meteor Observatory – not proper - XML

SKYBOT – Minor planets ephemerides (1840-2019)

Interest of climatology, meteorology

New branch of Science = e-Science

Theory VO (TVO)

- Methods of VO (parameters in DB, SQL...) for study of results of simulations , catalogues of simulated objects like SDSS...(PCA)
- Browsing of simulation space along different axes – parameters, regions...
- Evolutionary tracks, Photo Dissociation Regions
- Formation of artificial galaxies, clusters – N body models (Millenium Run 10 billions, 25TB)
- Theoretical Spectra (GAVO – Rauch, GRID)

CIELO VO - line catalogue SLAP

SLAP Viewer Copyright ESAC, Spain

Server Selector

- SLAP Services
 - IASD
 - LERMA
 - NIST ATOMIC SPECTRA
 - CIELO SLAP
 - <http://esav02:8080/cieloslapToolKit/cieloslap.jsp?>

Molecular line databases

Range of Search (m)

Wavelength Start Wavelength End

Select

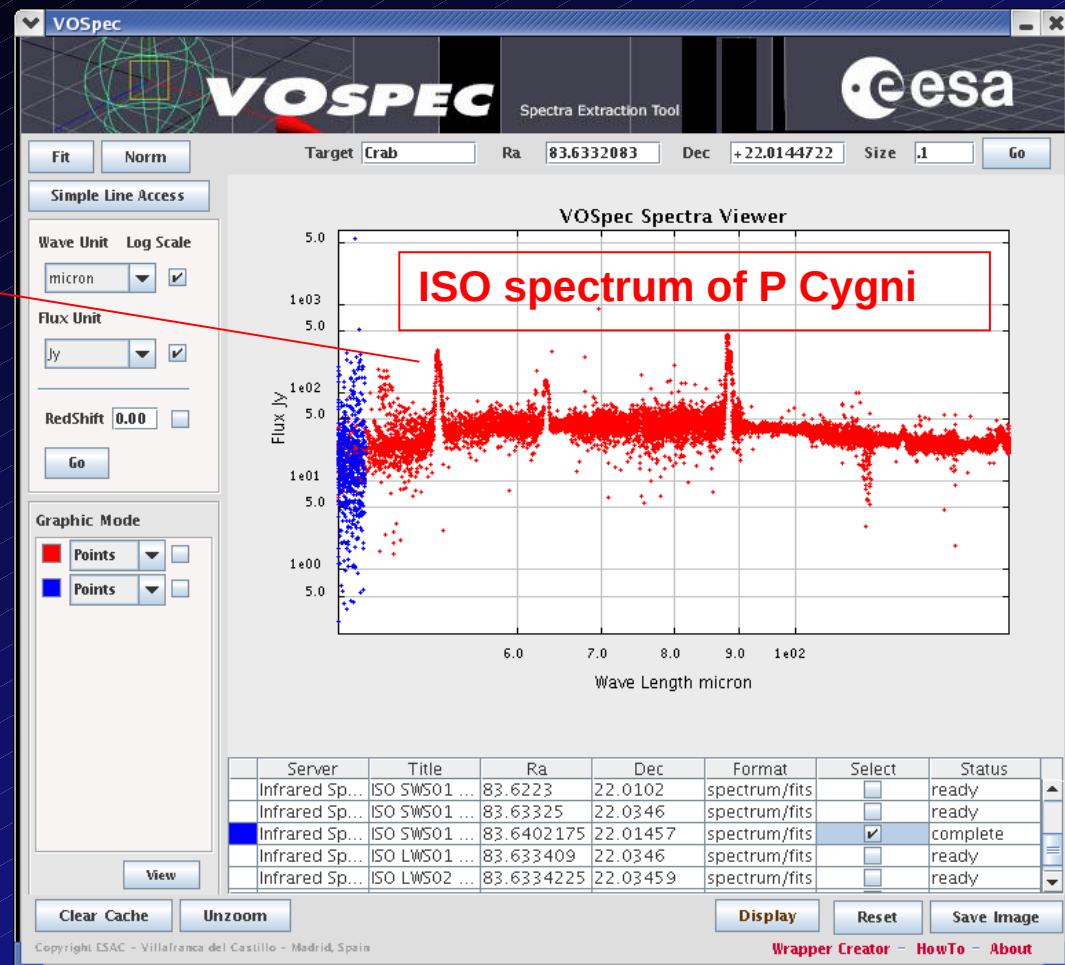
Reset

Slap Services Output

CIELO SLAP

Idm:Line.wavelength	Idm:Source...	Source.co...	Source.co...	Idm:Li...	Idm:...	Id...	Id...	Idm:...	Id...	
1.8627e-09	NGC1068	40.66963	-0.01328	1s_3p	1s2	1P1	150	OVII
1.7768e-09	NGC1068	40.66963	-0.01328	1s_4p	1s2	1P1	150	OVII
1.89671e-09	NGC1068	40.66963	-0.01328	2p	1s	2...	2...	OVIII
2.47793e-09	NGC1068	40.66963	-0.01328	2p	1s	2...	2...	NVII
2.21012e-09	NGC1068	40.66963	-0.01328	1s_2s	1s2	3S1	150	OVII
2.1602e-09	NGC1068	40.66963	-0.01328	1s_2p	1s2	1P1	150	OVII
2.18071e-09	NGC1068	40.66963	-0.01328	1s_2p	1s2	3P1	150	OVII
2.16210e-09	NGC1068	40.66962	0.01279	1s_2p	1s2	3P1	150	OVII

Close



VO for Atomic and Molecular Data

VAMDC (06/2009-12/2012 FP7)

13 organizations

Virtual Atomic and Molecular Data Centre

VO principles (web services, integration, registry,
SAMP, VODesktop, TOPCAT, VOSSpec)

(includes VALD extractor, NIST)

extended citation system (all providers acknowledged)

Access protocols in VO: TSAP

Theoretical models in the VO

• Theoretical spectra: TSAP

- Included in the SSAP standard (use case for theoretical spectra)
- A simple protocol.
- Dialog server-application.

The screenshot shows the SVO TSAP Interface web page. At the top, there is a navigation bar with links for 'Theoretical model services', 'Documents', 'Models', 'Services', and 'Funded by INTA'. Below the navigation bar, there is a logo for 'SVO Spanish Virtual Observatory' and a link to 'MINISTERIO DE CIENCIA E INNOVACIÓN'. The main content area is titled 'TSAP Interface' and describes it as 'An interface to test TSAP services'. It lists 'Services: VOSA Filters TSAP S3if' and provides email contact information 'esm@laeff.inta.es'. On the right side of the main content area, there is a 'Logout' link. Below the main title, there is a section titled 'TSAP Interface' with a brief description of the service: 'SVO Theoretical Data Access Service: ATLAS9 Kurucz ODFNEW/NOVER models (Castelli et al., 1997, AA, 318, 841)'. This section contains several input fields for specifying search parameters: 'teff_min:' (3500), 'teff_max:' (3500), 'logg_min:' (0.00), 'logg_max:' (0.00), 'meta_min:' (-2.50), and 'meta_max:' (-2.50). A 'Search' button is located at the bottom of this form. At the very bottom of the page, there is a link 'See metadata VOTable'.

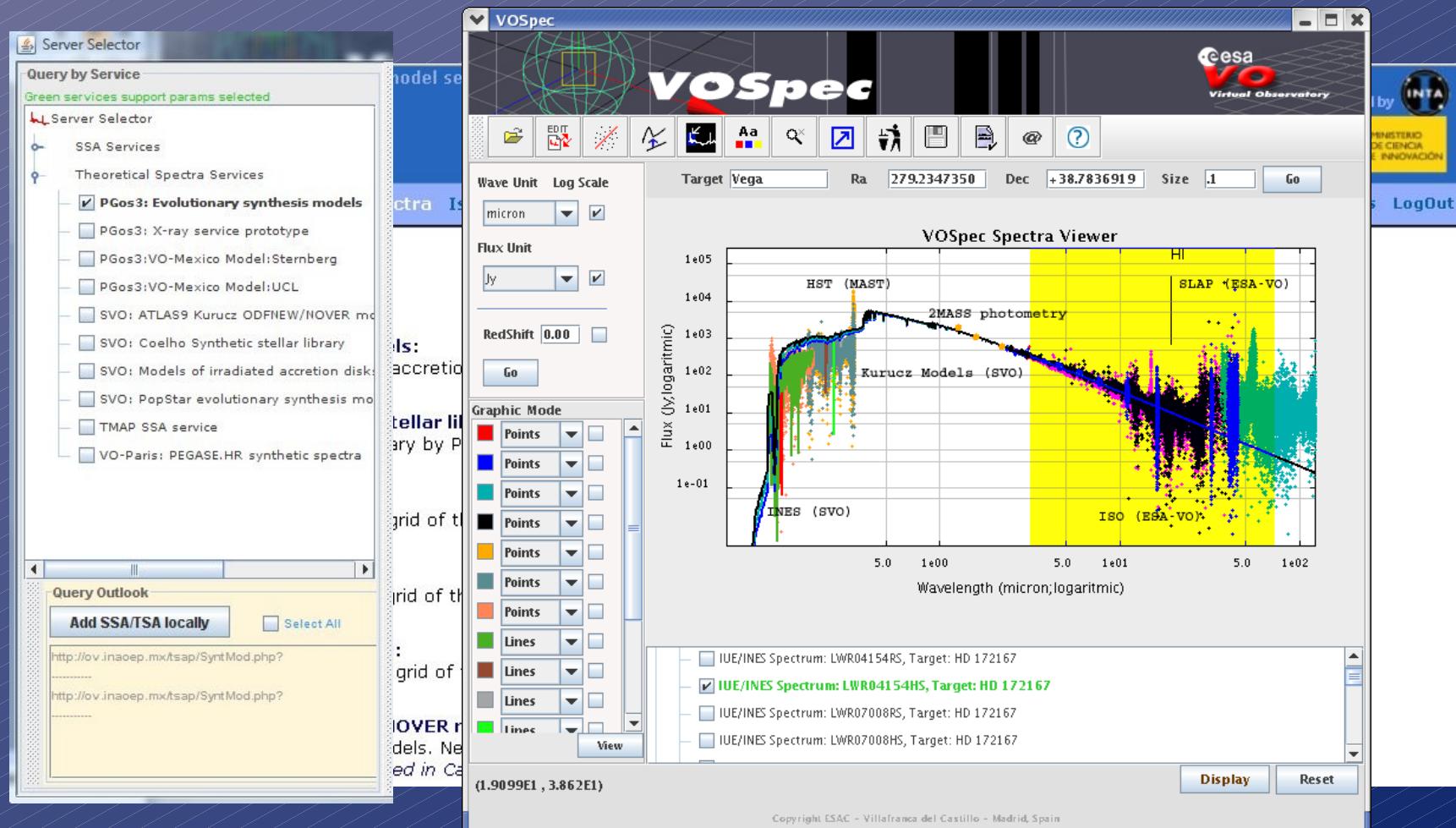
VOSpec - models by TSAP

The figure shows the VOSpec Spectral Analysis Tool interface. The top left panel displays a list of services, with several checked (e.g., The ISO Data Archive InterOperability System, VVDS-F02 DEEP spectra, Wisconsin Halfwave Spectropolarimeter, Wisconsin Ultraviolet Photo-Polarimeter Experiment, cutout server of HEROS archive of Ondrejov observations). The top middle panel is a 'Query by params' window showing a tree structure with a selected query for target 'p cyg' at position POS 304.44667416667, 38.03293027778 with size 0.1. The bottom left panel shows a 'Query Outlook' with URLs for various services. The main central area features a plot titled 'VOSpec Spectral Analysis Tool' showing the spectrum of HD 141569. The x-axis is 'Wavelength (micron; logarithmic)' ranging from 1 to 50. The y-axis ranges from -14 to -06. The spectrum is shown as a series of colored points (red, green, orange) with a blue line representing the model fit. Below the plot are various analysis tools and a legend for data types. The top right corner shows the 'cesa VO Virtual Observatory' logo.

Archives, Theory, VO-Science, DataMining, E&O

Simple Spectral Access Protocol V1.04

Appendix A: Theoretical Spectral Access Use Case



- Other VO Data Centres providing theoretical spectra using TSAP

**GERMAN ASTROPHYSICAL
GAVO
VIRTUAL OBSERVATORY**

German Astrophysical Virtual Observatory

Archive: **TMAP Spectra** [More information on archive](#)

Effective temperature in K: +/-

Surface gravity (log g) in cm/s²: +/-

Mass fraction 0:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 1:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 2:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 3:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 4:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 5:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 6:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>
Mass fraction 7:	<input type="button" value="none"/>	<input type="text" value="0.0"/>	<input type="text" value="+/- 0.3"/>	<input type="text" value="dex"/>

Band: The wavelength range in format "wavelength₁/wavelength₂" in the selected unit.

Data format: Format of the individual spectra. (No need to select, if return format is html.)

Return Format: votable html The format in which to present the metadata. (If html is selected, no further selection of data format is necessary, since links to all available formats will be created anyways.)

- PGos3 (Mexico), PEGASE (VO-Paris)

BaSTI database



Micro-simulations inside the VO: the BaSTI case

EURO^{VO}



P. Manzato⁽¹⁾, M. Molinaro⁽¹⁾, F. Gasparo⁽¹⁾, F. Pasian⁽¹⁾, A. Pietrinferni⁽²⁾, S. Cassisi⁽²⁾, C. Rodrigo⁽³⁾, M. Cerviño⁽⁴⁾, E. Solano⁽³⁾
INAF - SI / Trieste Astronomical Observatory; (2) INAF – Teramo Astronomical Observatory; (3) LAEFF-INTA / Spanish VO; (4) Instituto de Astrofísica de Andalucía – CSIC / Spanish VO

S3P (Simple Self-described Service Protocol) implementations

In collaboration with SVO (the Spanish Virtual Observatory) we presented S3P in the last IVOA Interoperability Meeting. S3P (Simple, Self-described Service) is a protocol oriented to handle theoretical data in the VO framework. It is based in the ability of the data server to describe itself in a simple standardized way.

This is a step by step protocol:

1 step: the service described it self (input and output parameters);

<http://myservice.com/s3.php?format=metadata>

2 step: http query and response in VOTable format;

<http://myservice.com/s3.php?param1=value1¶m2=value2...>

3 step: retrieve the simulated files of interest via http GET;

<http://myservice.com/s3.php?id=12>

We developed two prototype implementations of S3P for BaSTI: one for isochrones and one for tracks:

<http://albione.oa-teramo.inaf.it/PHPmetadata/BaSTIisochron.php?format=metadata>

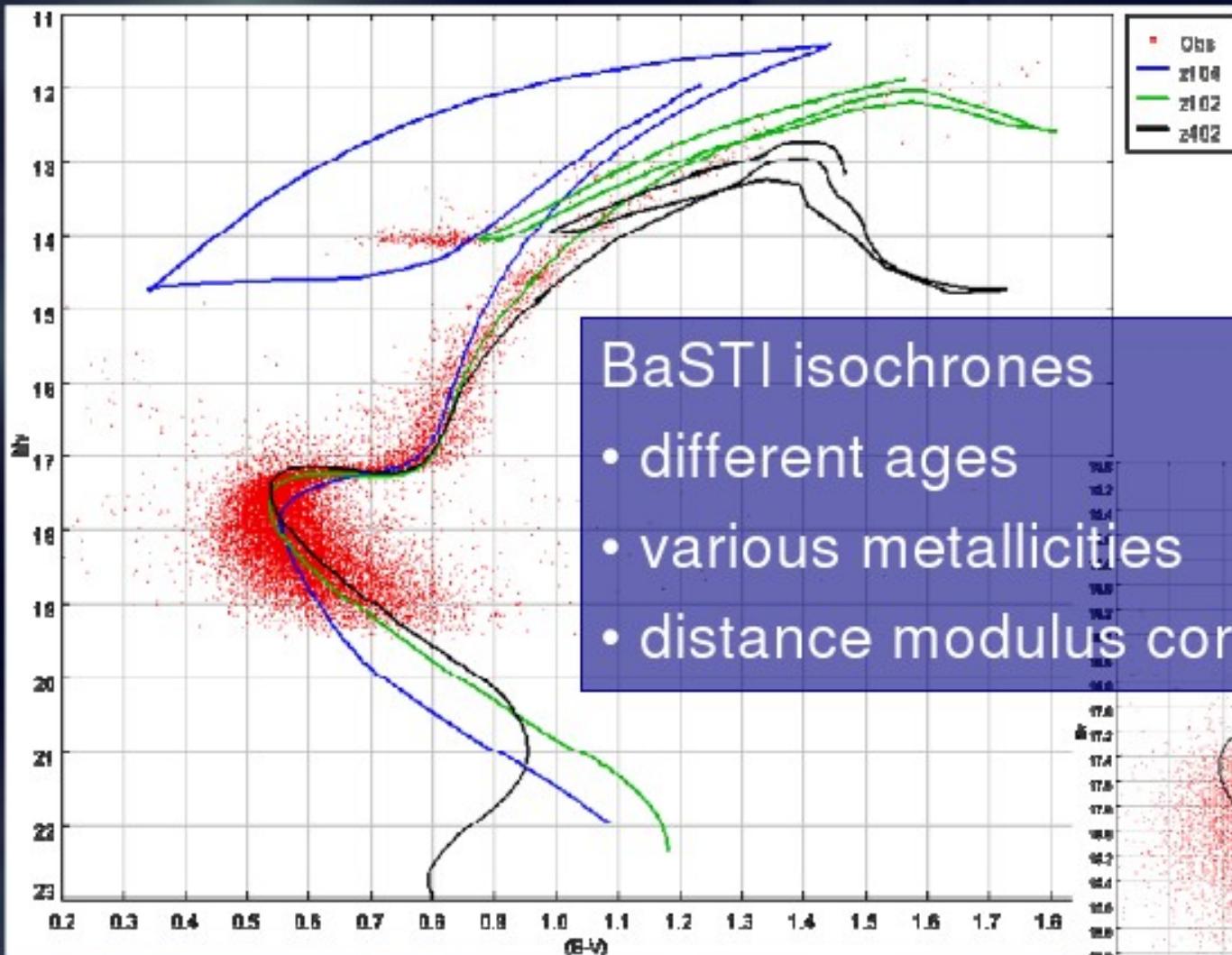
<http://albione.oa-teramo.inaf.it/PHPmetadata/BaSTItrack.php?format=metadata>

Param	UCD	Description
INPUT:age_min	time.age	Min. age of the isochron in Gyr (min value 0.03 Gyr)
INPUT:age_max	time.age	Max. age of the isochron in Gyr (max value 19 Gyr)
INPUT:meta_min	phys.abund.Z	Min. mass fraction of the initial heavy elements abundance for stellar isochron model (min value 0.0001)
INPUT:meta_max	phys.abund.Z	Max. mass fraction of the initial heavy elements abundance for stellar isochron model (max value 0.4)
OUTPUT:age	time.age	value for the stellar Age for the model. Age is given in Gyr
OUTPUT:meta	phys.abund.Z	value of mass fraction of the initial heavy elements abundance for the model.
OUTPUT:[M/H]	phys.abund.Z	The metal abundance in the spectroscopic formalism.
OUTPUT:[Fe/H]	phys.abund.Fe	The iron abundance in the spectroscopic formalism.
OUTPUT:Y	phys.abund.T	value of mass fraction of the initial helium abundance. Actually calculated as Y = 1.44*(Z-0.0001).
OUTPUT:MassLoss	phys.mass.loss	value of mass loss according to the Reimers (1975) law.
OUTPUT:title	VOX.Image_Title	Title.



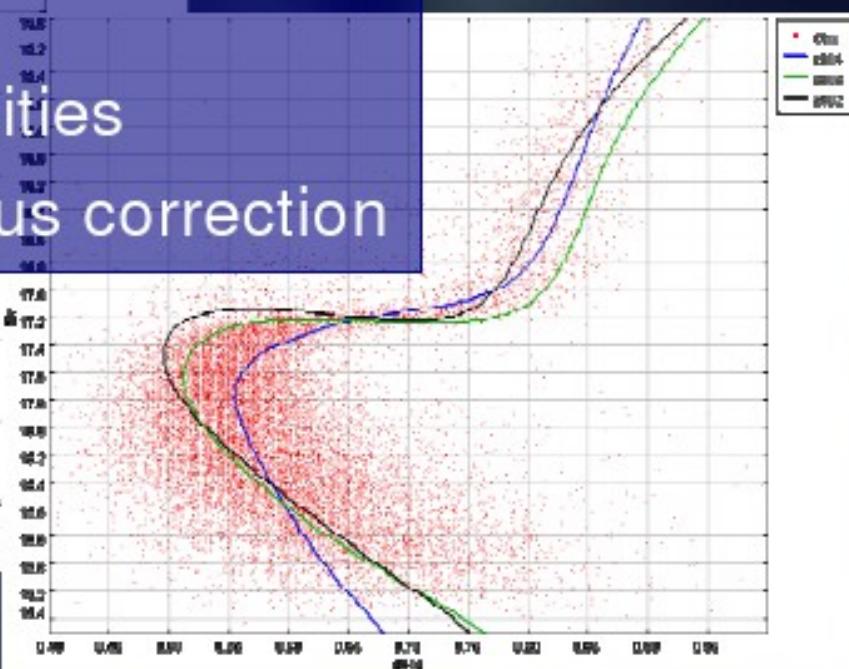
BaSTI Isochrones

step 1
metallicity



BaSTI isochrones

- different ages
- various metallicities
- distance modulus correction



$$z = 0.01 \text{ (\alpha-enh)} ; 0.008 \text{ (scaled solar)}$$

Archives, Theory, VO-Science, DataMining, E&O

Theoretical model services

VOSA: VO Sed Analyzer
VO SED Analyzer

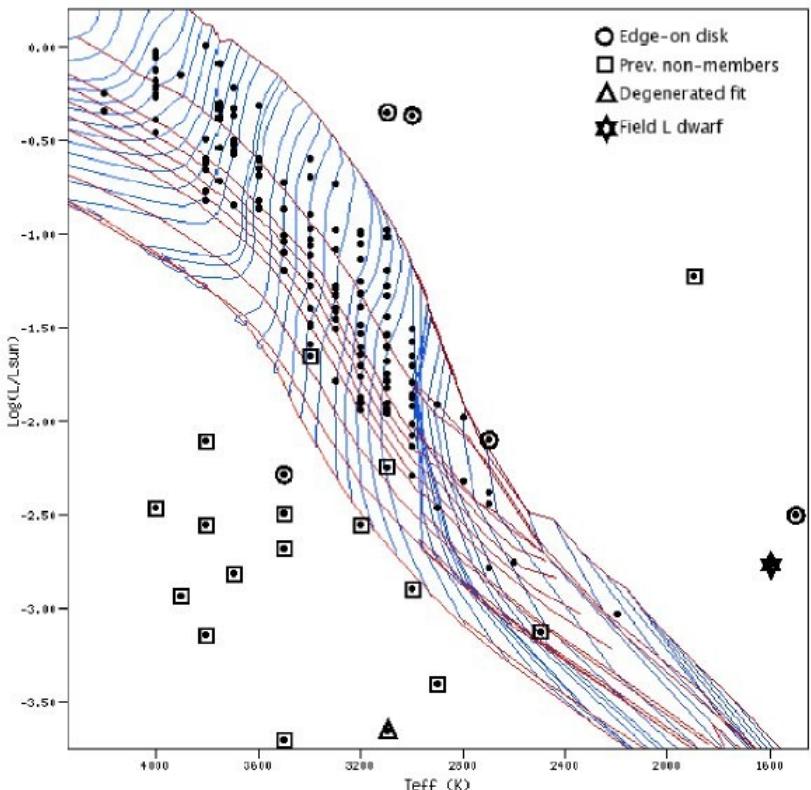
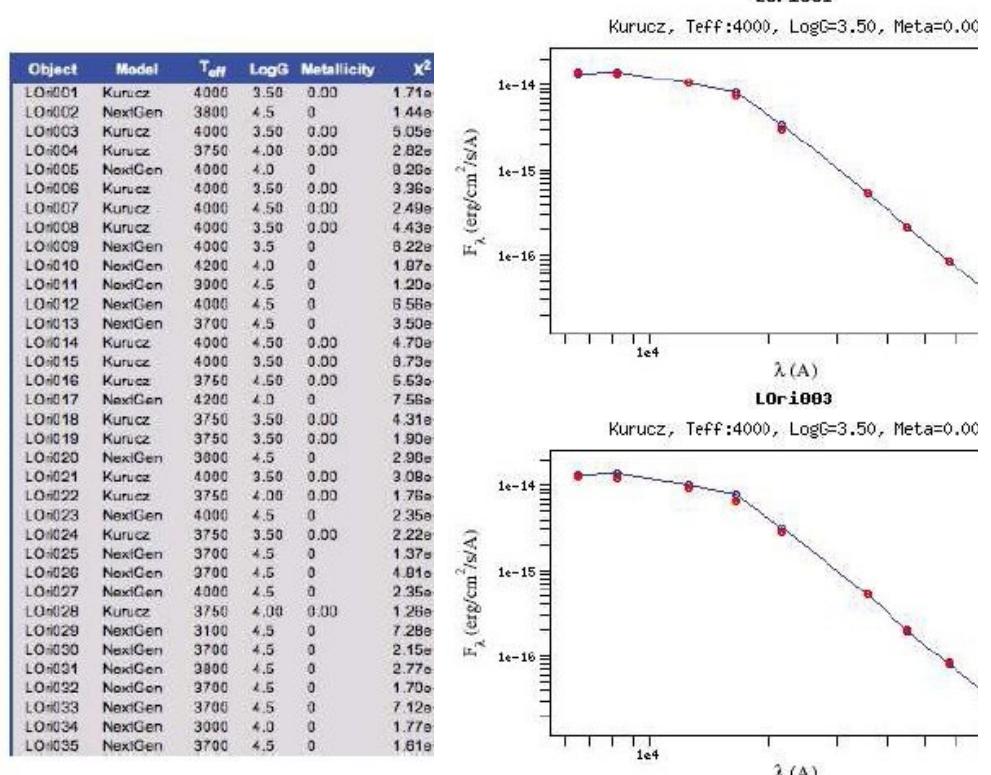
Services: VOSA Filters TSAP S3if

Astronomy & Astrophysics manuscript no. Synth'VO'PRI'ref format
August 2, 2008

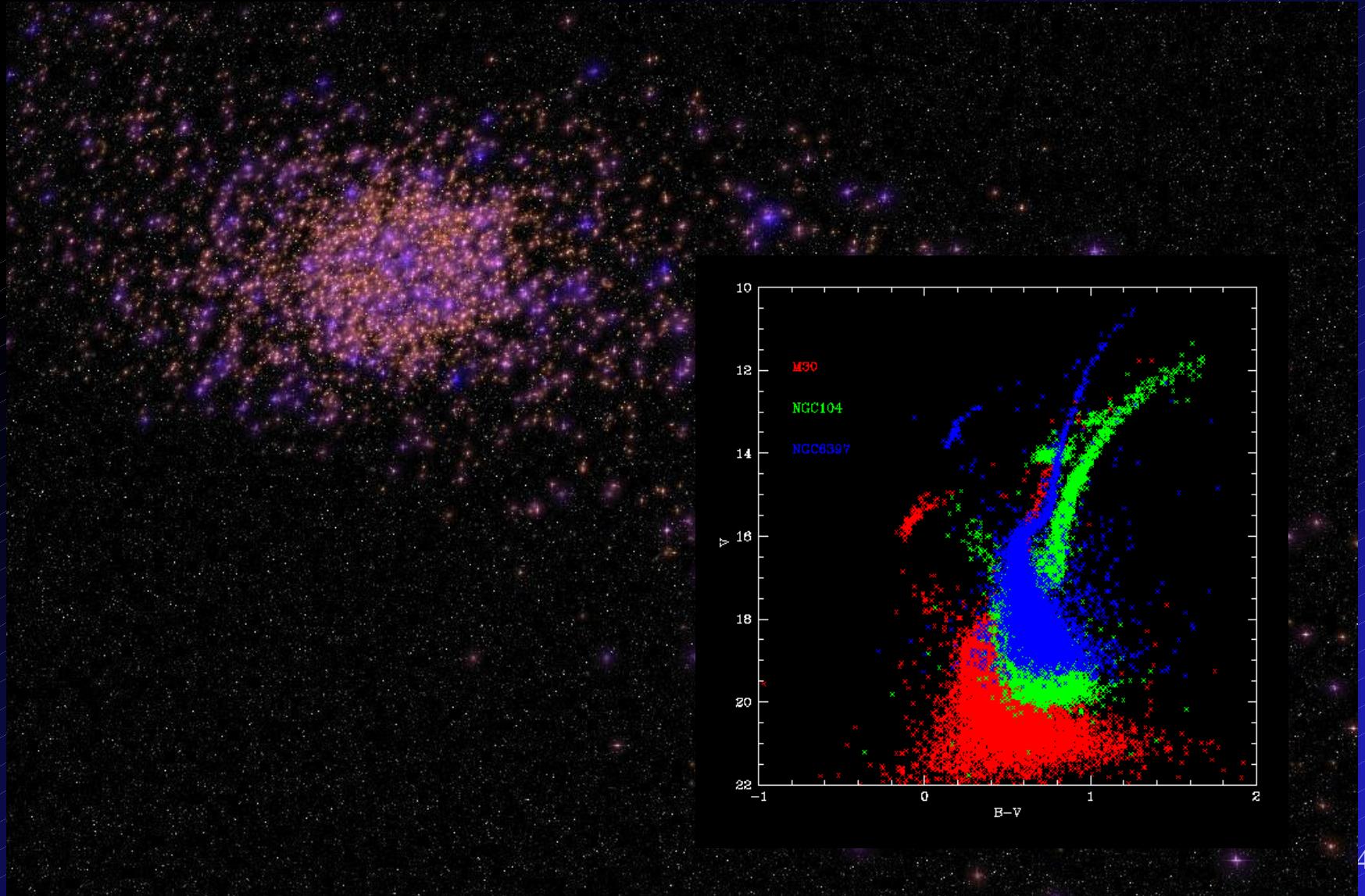
VOSA: Virtual Observatory SED Analyzer.

An application to the Collinder 69 open cluster

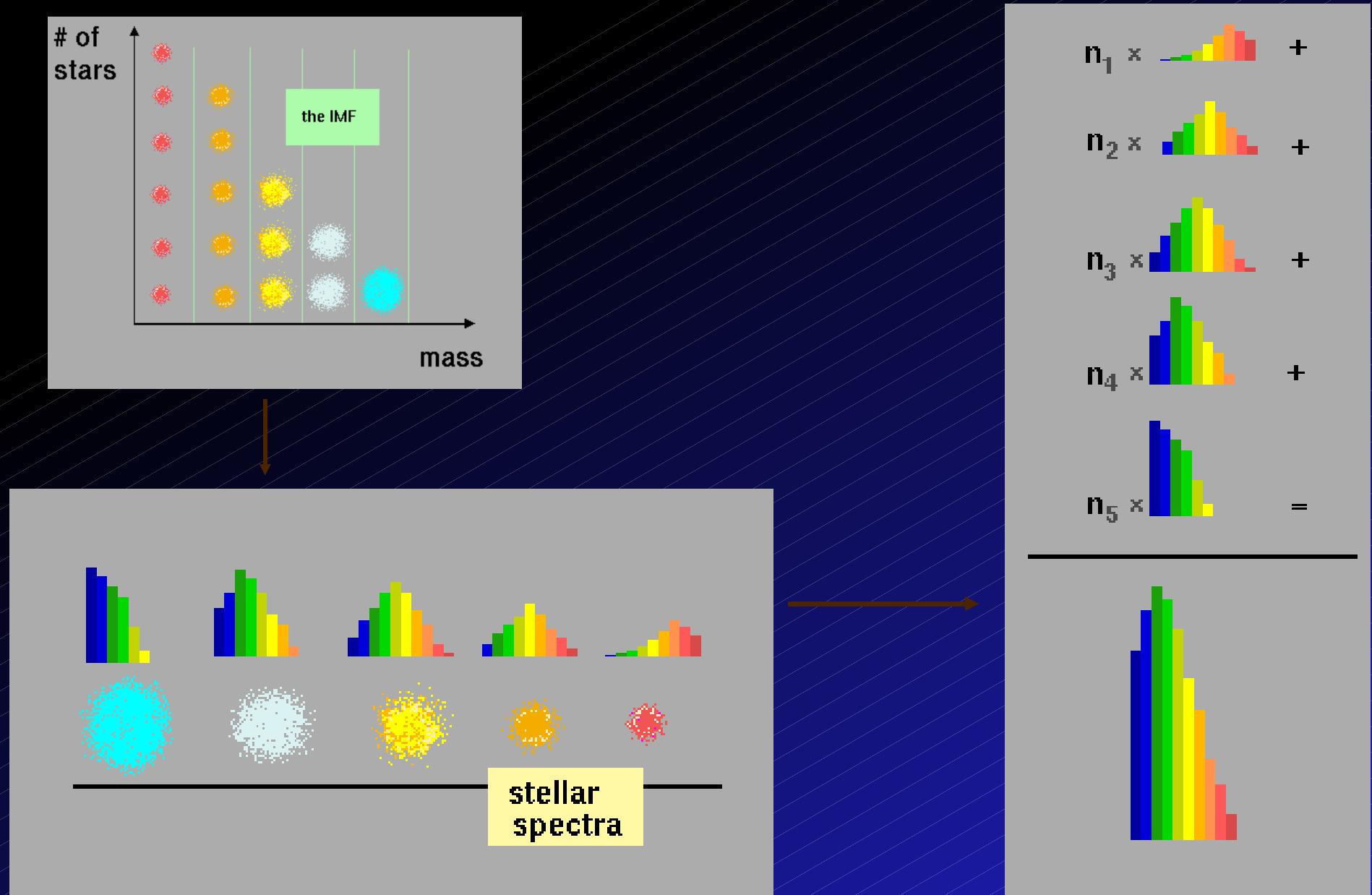
A. Bayo^{1,2}, C. Rodrigo^{1,2}, D. Barrado y Navascués^{1,2}, E. Solano^{1,2}, R. Gutiérrez^{1,2}, M.



N Body Simulations of Globular Cluster Evolution



Stellar populations are modeled with synthesis models



(Available as *Theoretical Simple Access Protocol* server :<http://ov.inaoep.mx>)

Using SimDB/SimDAP

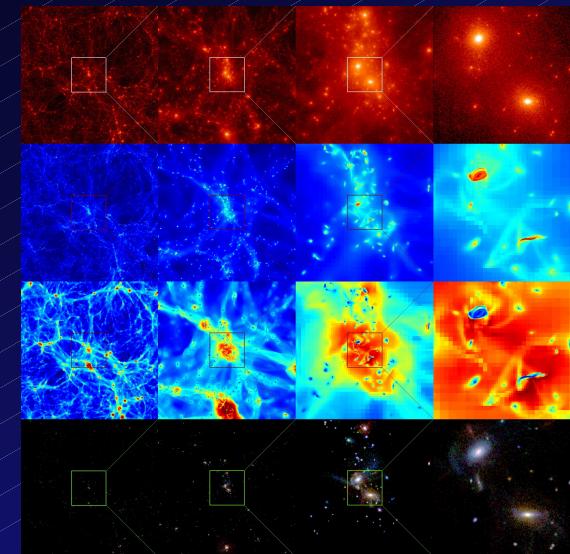
- Cosmological simulations
 - Prototypes for GalMER, Horizon
- PDR simulations
 - test implementation of Meudon PDR code
- Isochrones/evolutionary tracks
 - BaTSI
- Visualization tools
 - VisIVO

GalMer

DB Query Query Results Experiment Snapshot Description

Select Input Parameters

Galaxy #1	Galaxy #2	Query
gE0 ▲ gSa gSb gSd ▾	gE0 ▲ gSa gSb gSd ▾	Orbit type 1 ▾ Spin Prograde ▾ Inclination 0 deg ▾



Virgo - Millennium Database

Documentation CREDITS/Acknowledgments Registration News Databases millenium (context) VIRGO GAVO

Check out the latest news about the release of the Millennium-II database.

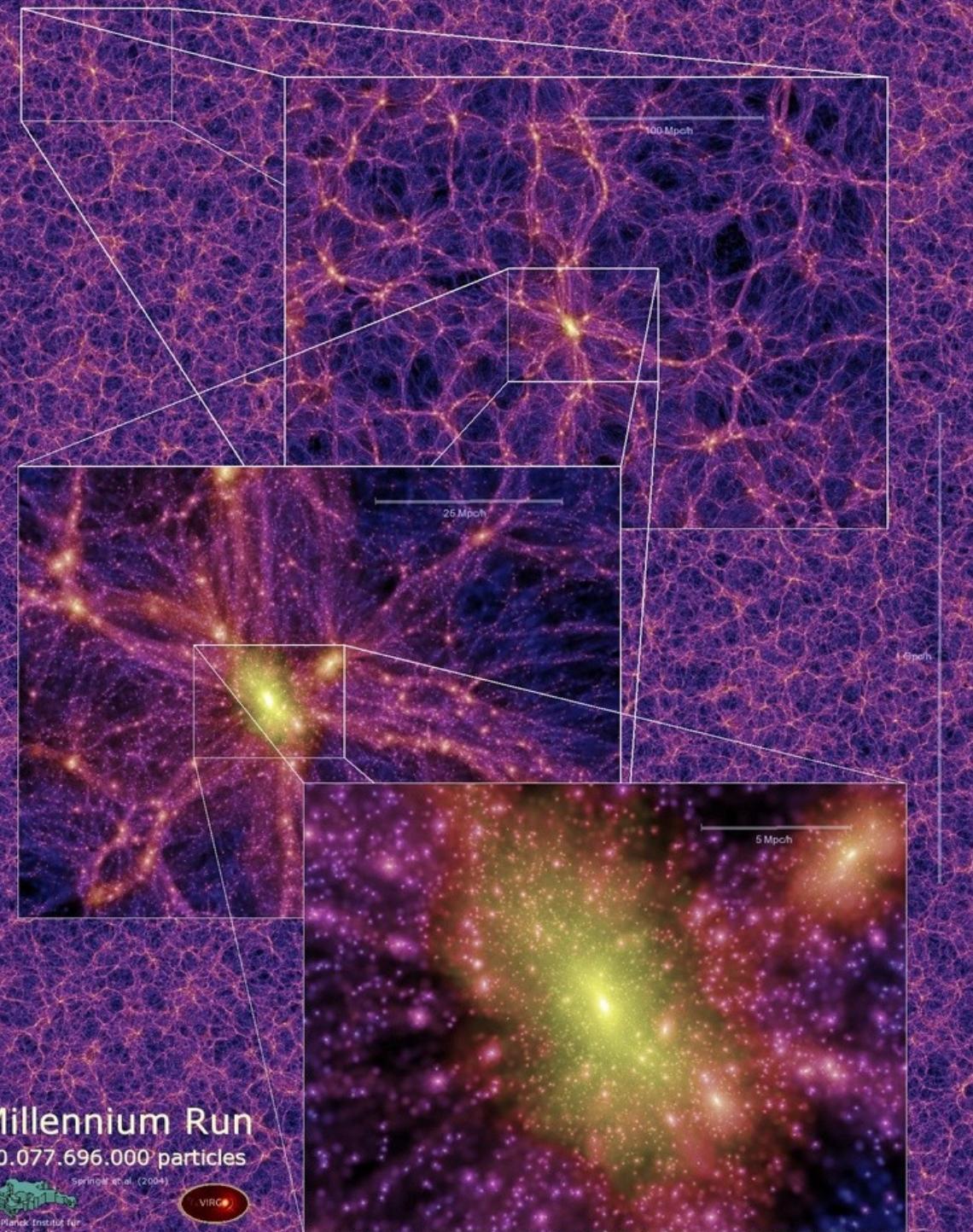
Streaming queries return unlimited number of rows in CSV format and are cancelled after 30 seconds. Browser queries return maximum of 1000 rows in HTML format and are cancelled after 30 seconds.

Query (stream)
Query (browser)
Help

Maximum number of rows to return to the query form: 10 ▾

GADGET-2: Galaxies with dark matter and gas interact

A code for cosmological simulations of structure formation



Millenium Run

10^{10} particles

Several Gpc to

10 kpc

Cube 2 billion ly

One month MPSSC

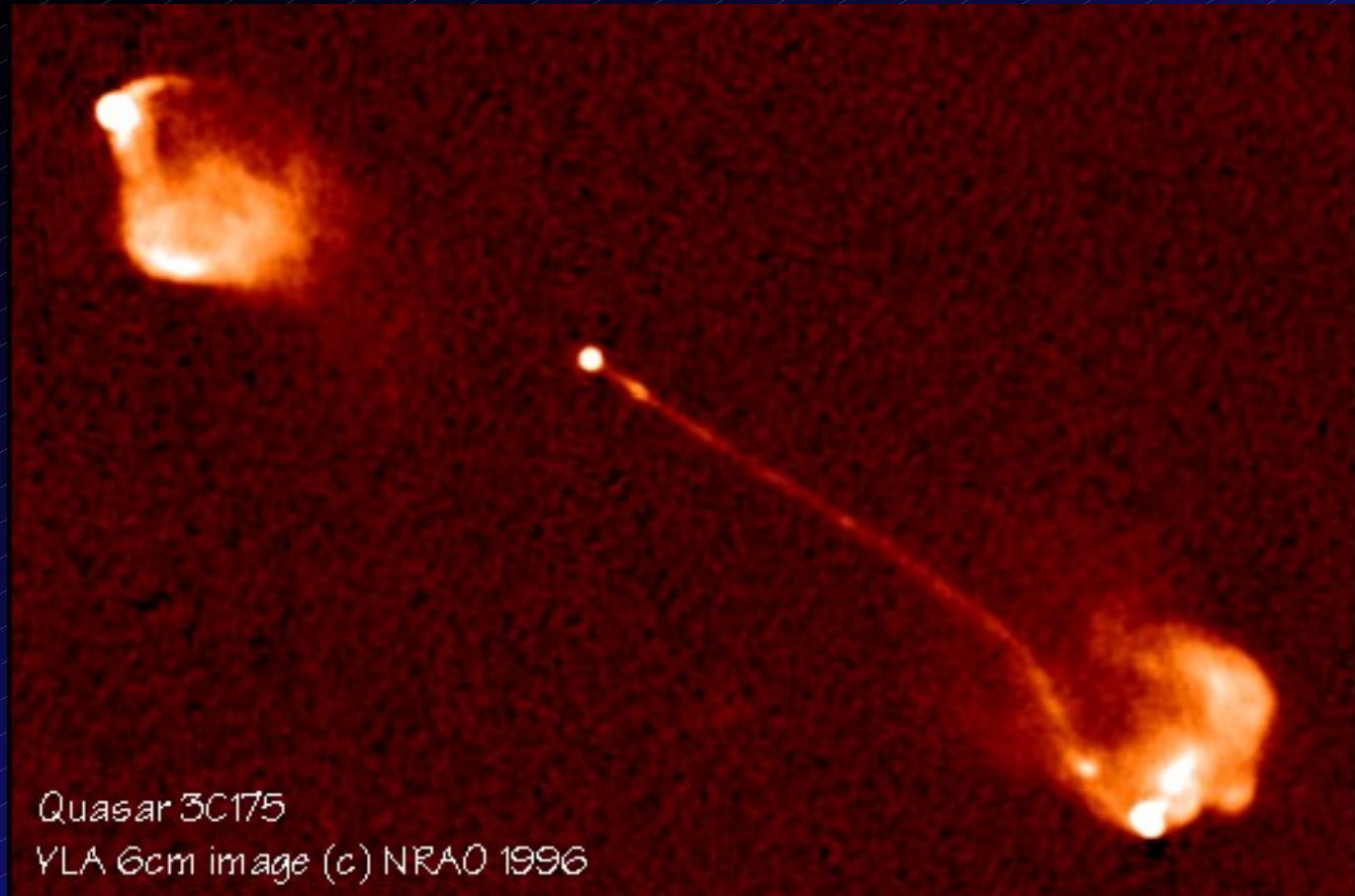
25 TB

Evolution of 20 mil
galaxies

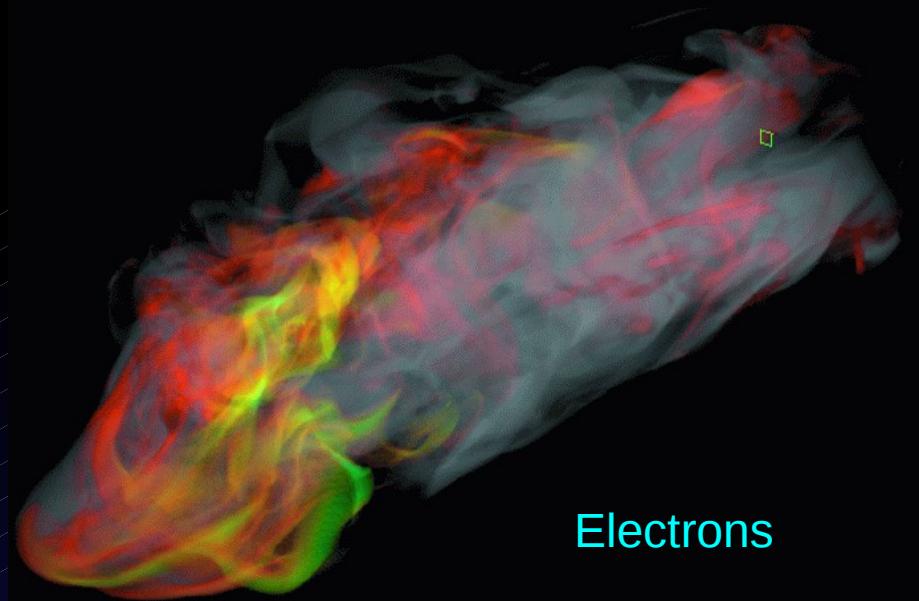
Evolution merger tree

Collimated Outflows from AGN

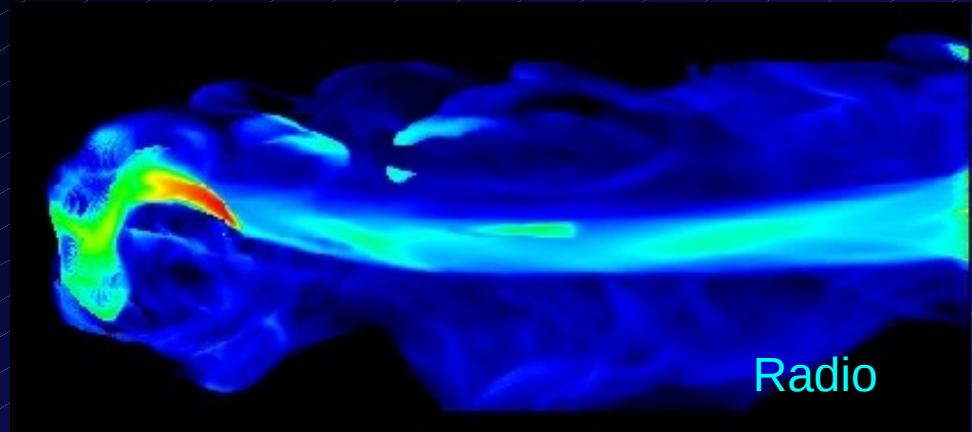
- 3C 175



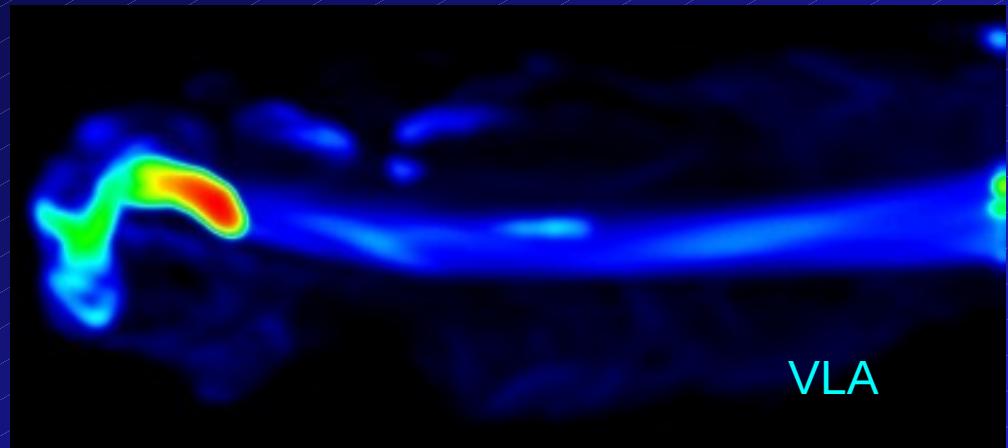
MHD Simulations of Collimated Outflows from AGN - Virtual Telescope Observations



Electrons



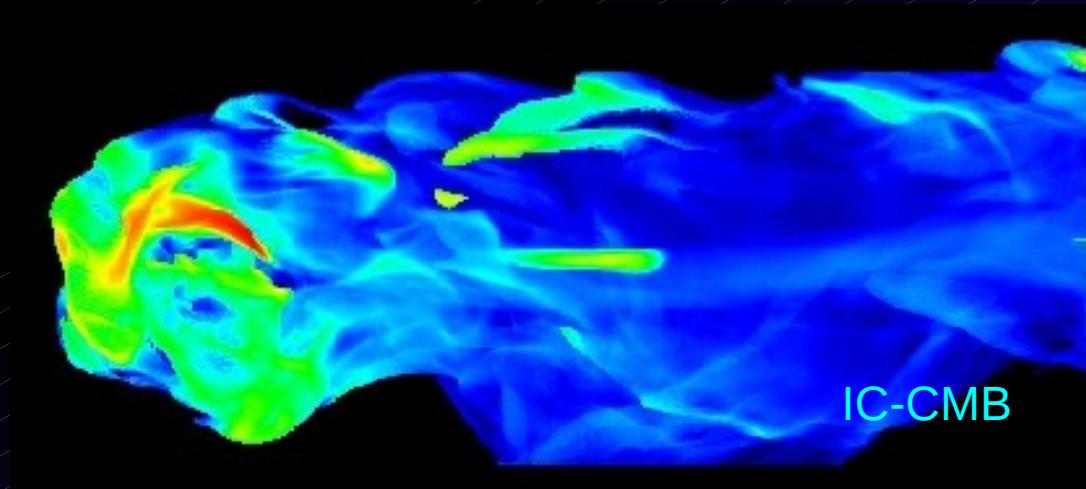
Radio



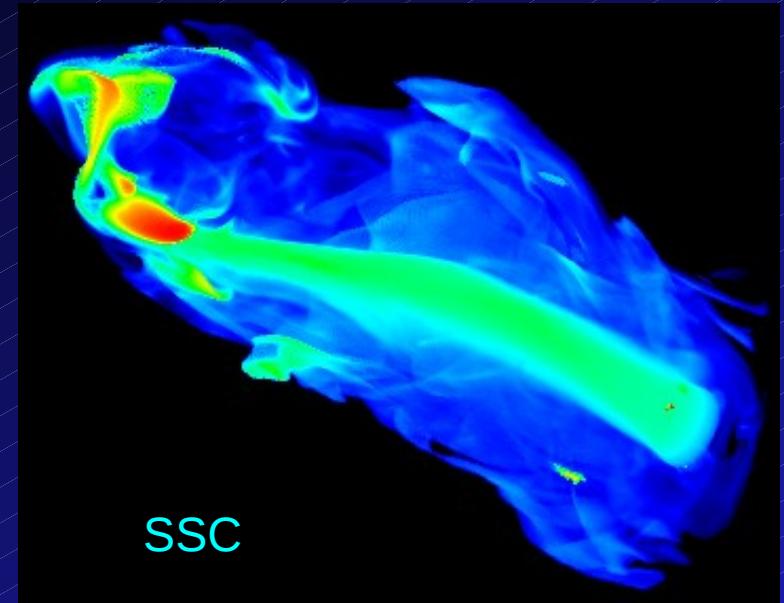
VLA

Compare with
Radio
Archives

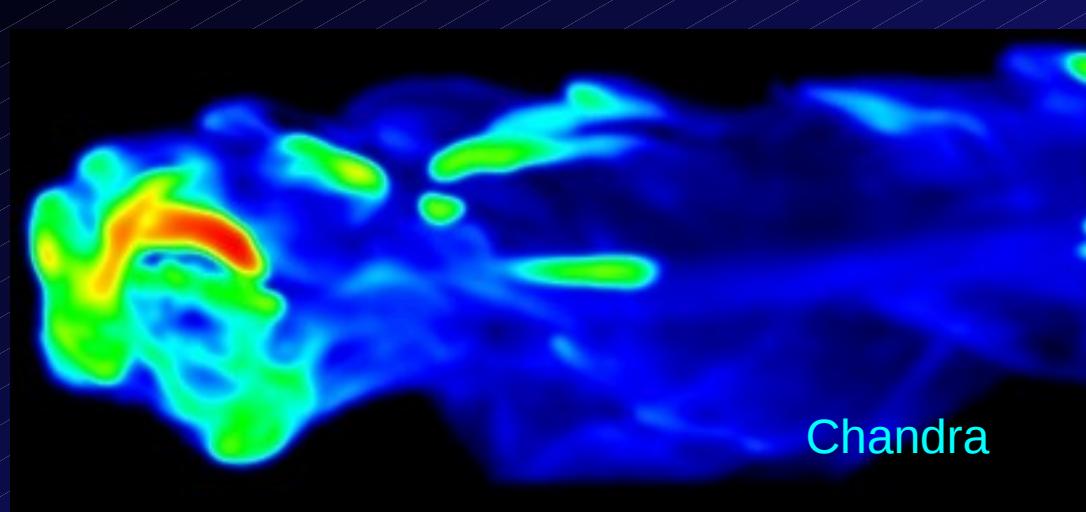
MHD Simulations of Collimated Outflows from AGN - Virtual Telescope Observations



IC-CMB



SSC

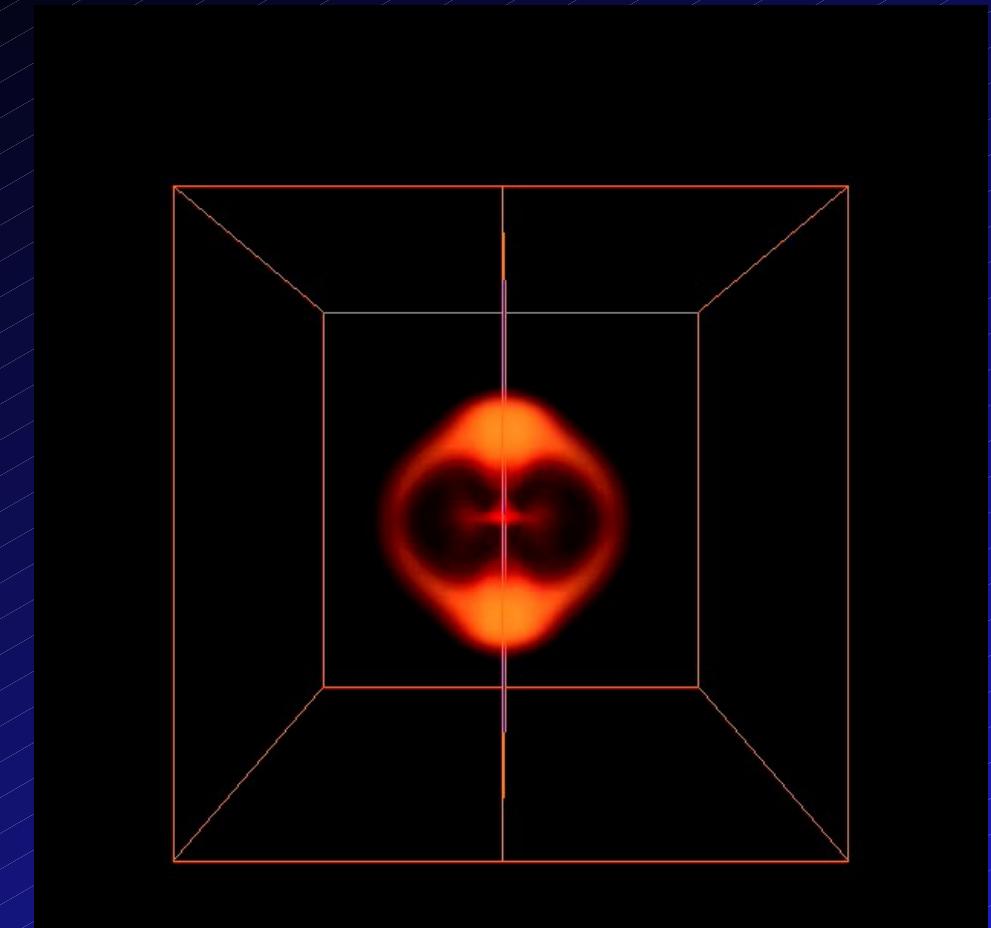
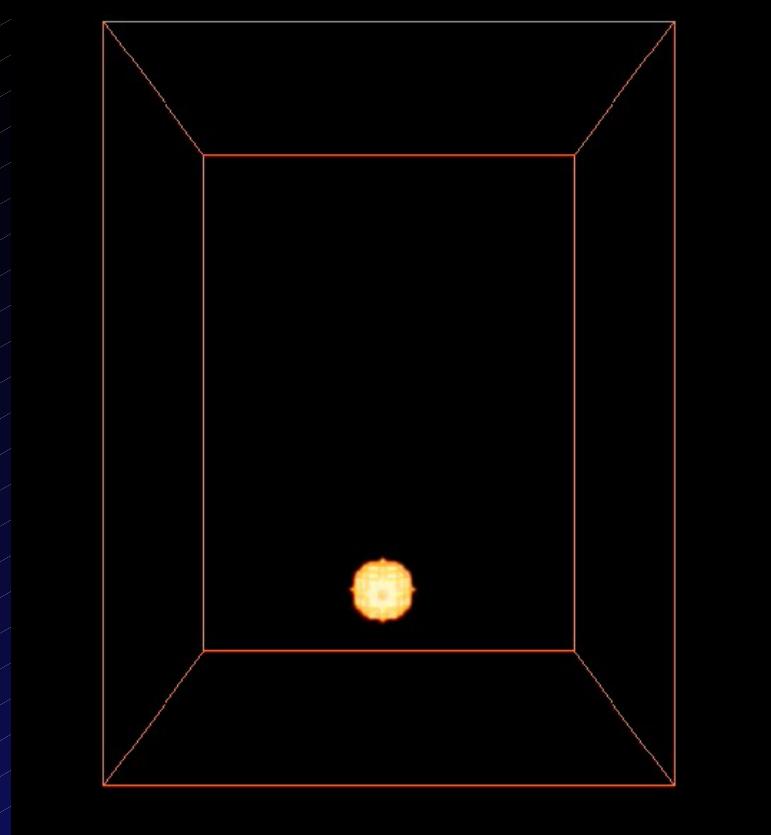


Chandra

Compare with
Chandra Archives

Three Dimensional MHD Calculations

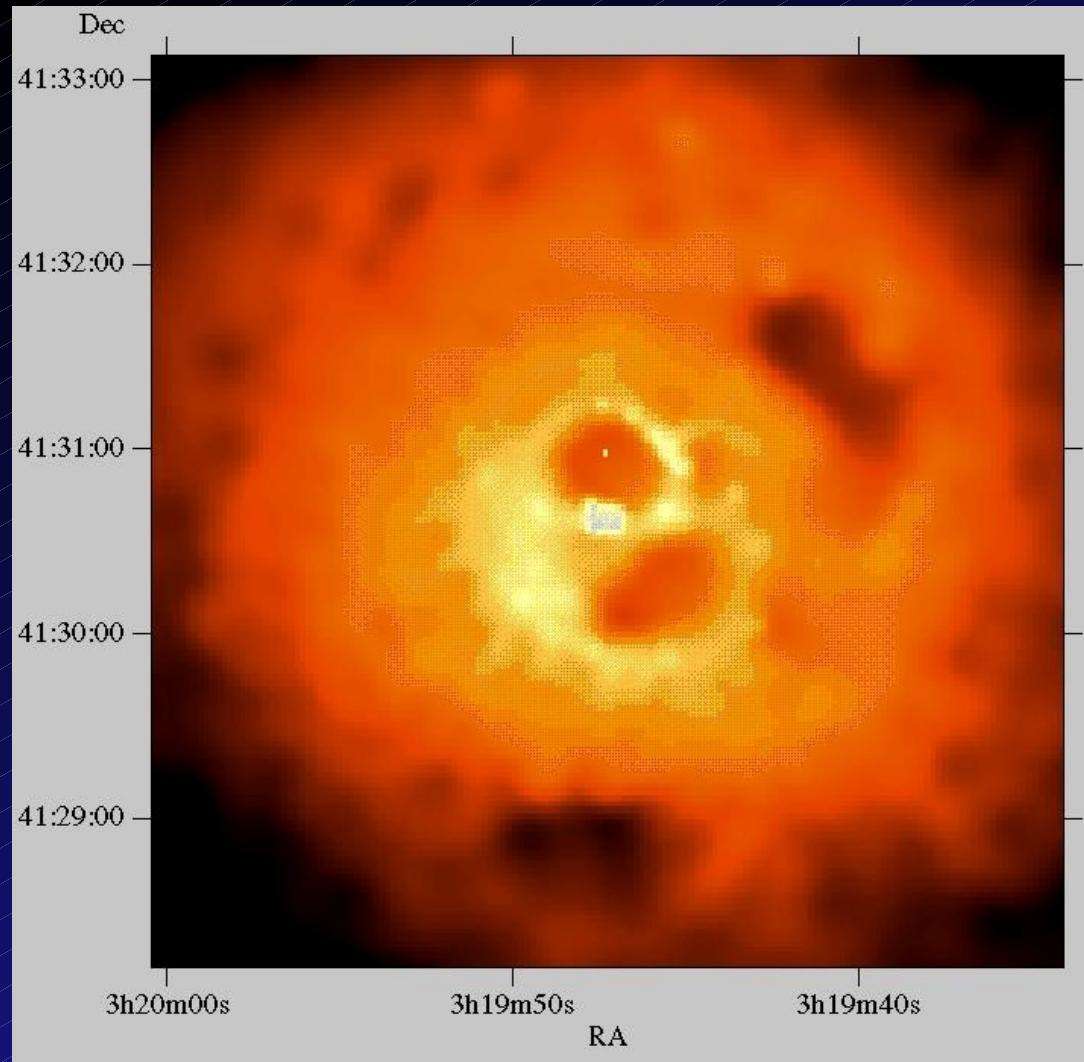
- $\beta = 3000$



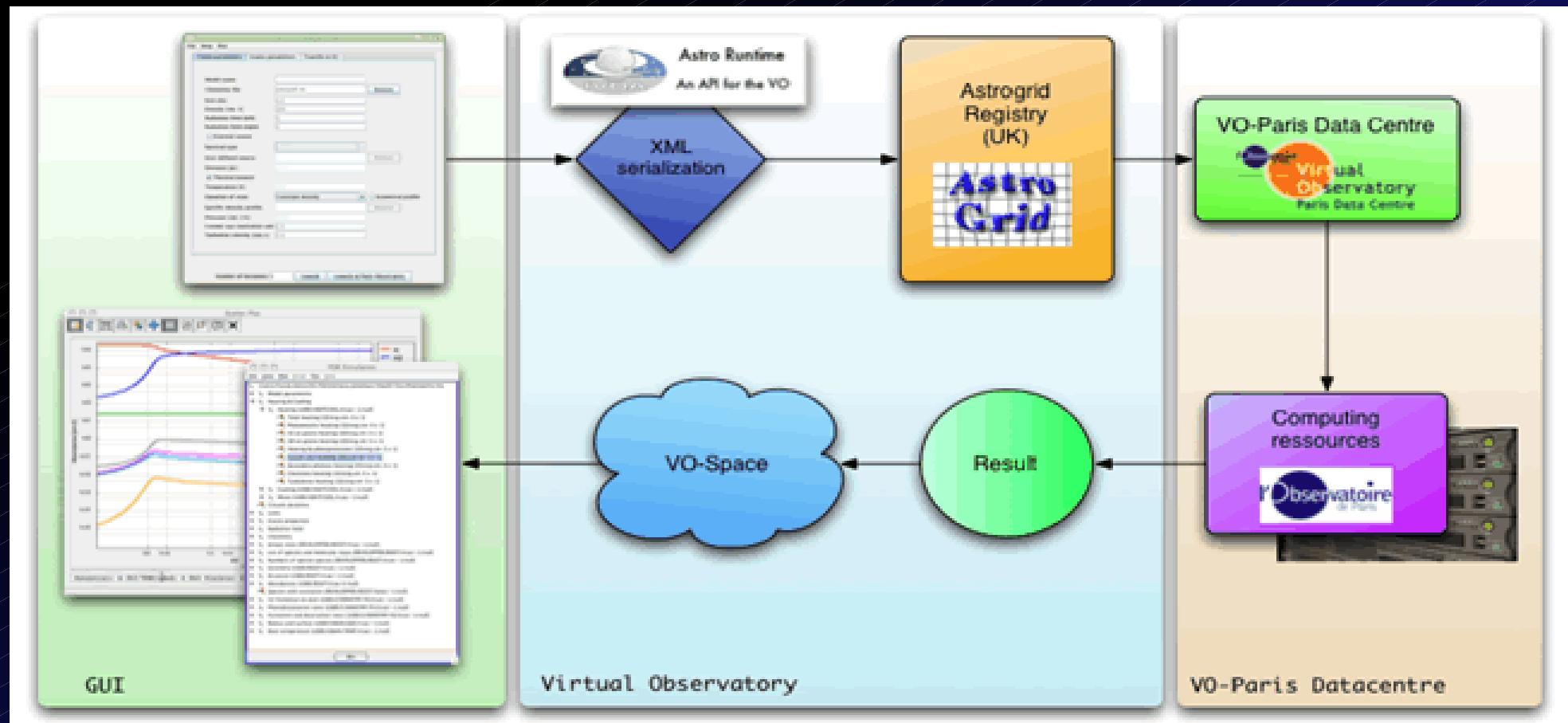
Relic Radio Bubbles in Galaxy Clusters

- N1275

Compare
with
Chandra
Archives



PDR VO-infrasctructure



PDR database and clients

PDR Database

Output Files

Code produces

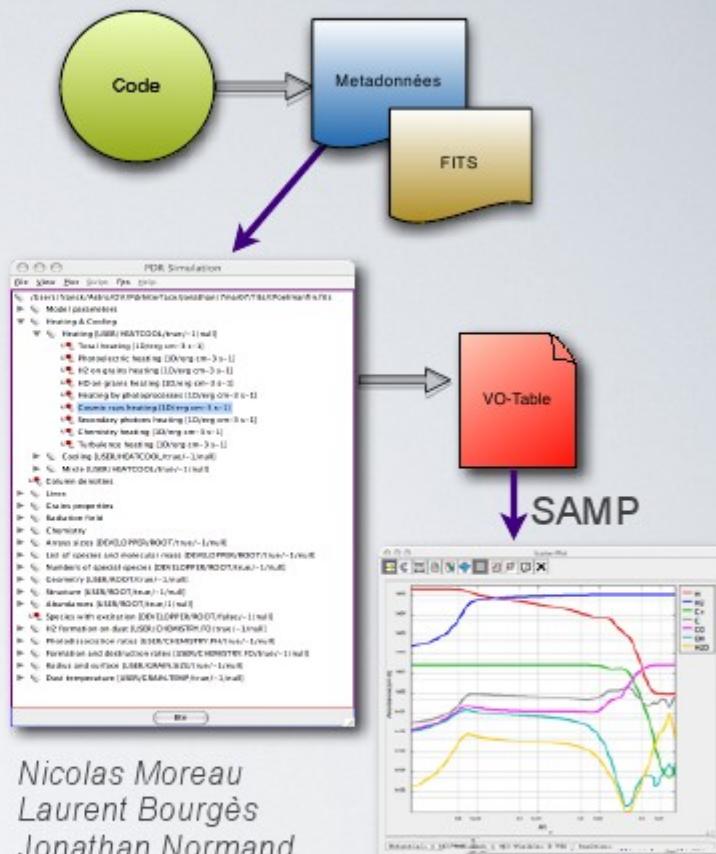
- raw data : FITS File
- XML / VO-TABLE : meta-data
(name, description, units, UCD, ...)

Provide all quantities computed by the code

- observables
- theoretical quantities

PDR Analyser

- browse the computed quantities
- extraction (ASCII, VO-Table)
- SAMP
- Download data from VO-Space
- Scriptable



PDR code via VODesktop

VO Explorer – PDR

Contents of PDR – 3 resources

Content – Subject: unknown, ???, chemistry, interstellar gas, interstellar matter, interstellar medium, interstellar molecules, models

Coverage – Waveband: unknown, infrared, millimeter, optical, radio, uv

Resource Type: CeaApplication, DataService, Organisation

Filter results:

Status	Flag...	Title	Capability	Date
●		Meudon PDR code	Meudon PDR code	2007-12-14
●		Meudon PDR code VO-Paris		2007-04-11
				2007-04-11

New Smart List

Actions: Execute Task

About: Selection: CeaApplication, Further Info, Email Curator

Information | Table Metadata

Meudon PDR code

Short Name: Meudon PDR code, ID: ivu/obspm.fr/pdr
Type: CeaApplication, Created: 1999-01-01T00:00:00, Updated: 2007-12-14T00:00:00

Content Type: other, Subject: ???
The Meudon PDR code is a tool to model the physics and the chemistry of interstellar gas at stationnary state. It considers a stationnary plan-parallel slab of gas and dust illuminated by a UV radiation field and solves radiative transfer, thermal balance and chemistry. It is then possible to deduce column densities and emissivities to compare to observations. The exact physics in the code is described on our website. [Further Information...](#)

This resource describes a Remote Application (CEA).
Interfaces: simple

Version 1.0, Dates: representative: 2006-01-12
Creator: [VO-Paris](#), creator logo: 

Annotate: Flag, Highlight, Alternative title, Notes, Tags, Monitoring service, No known providing services

-Paris Datacentre

Franck Le Petit
LUTH - Paris Observatory

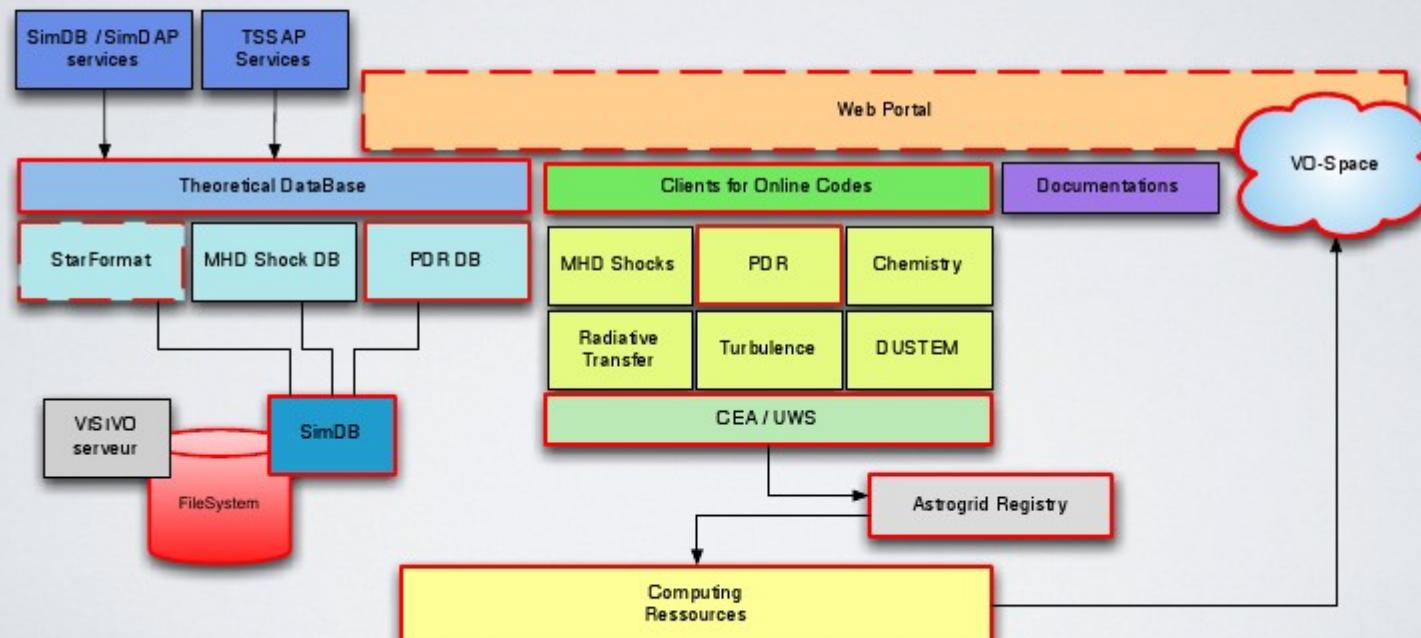
ISM platform

□ Interstellar Medium Platform

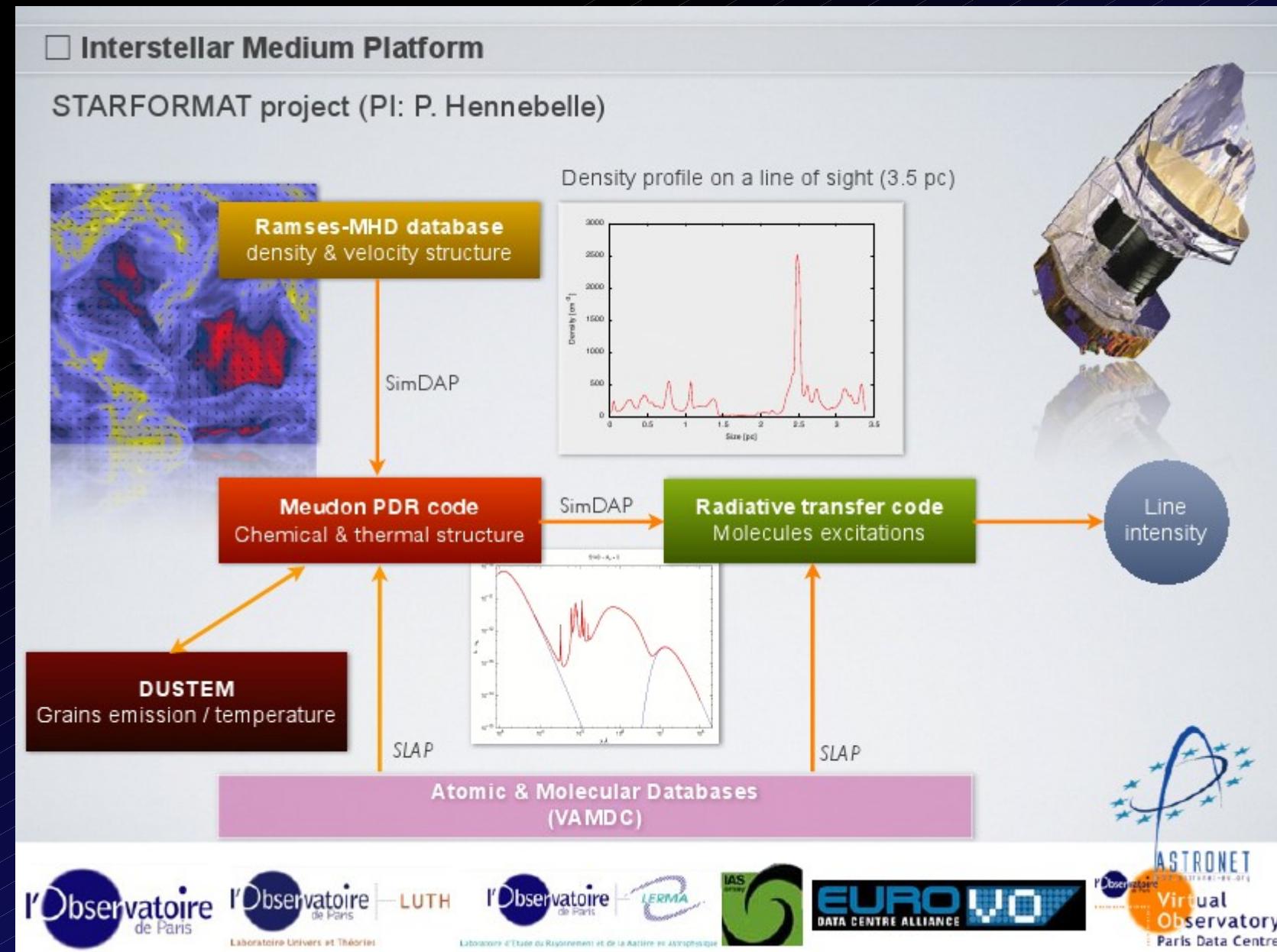
Bring together expertise in modeling / simulation of the ISM

Provide theoretical services about ISM

Codes - Databases - Tools & services



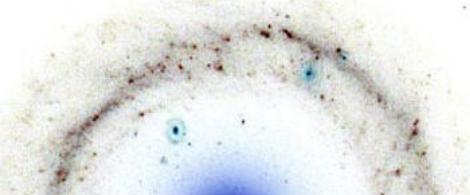
Complex join of TVO bricks



VO Science

- 31 (9) new obscured type 2 QSO (Padovani 2004)
- Brown dwarfs (about 20 candidates)
- Brightest (WD?) Albus-1 (Cabalero et al. 2008)
- Widest CPM binaries
- AGB to PNe - 100 new (200) with VO
- SED (Spectrum Energy Distribution)
- Bolometric magnitude
- VOEvent – robotic telescopes (GRB, transits,)
- Outreach , Education (MS WWT, GoogleSky)

BDs discovered using VO



PROJECT

Standards
Software & Services
Publications
Prototypes

Internal Logos

ABOUT NVO

What is the NVO?
Science Objectives

COMMUNITY

Discussion Lists
International VO
VOForum
Metadata (NCSA)
Other Links

PEOPLE

Contact Us
Personnel

Brown Dwarf Search Science Prototype: Real-Time Cross Matching of Large Catalogs

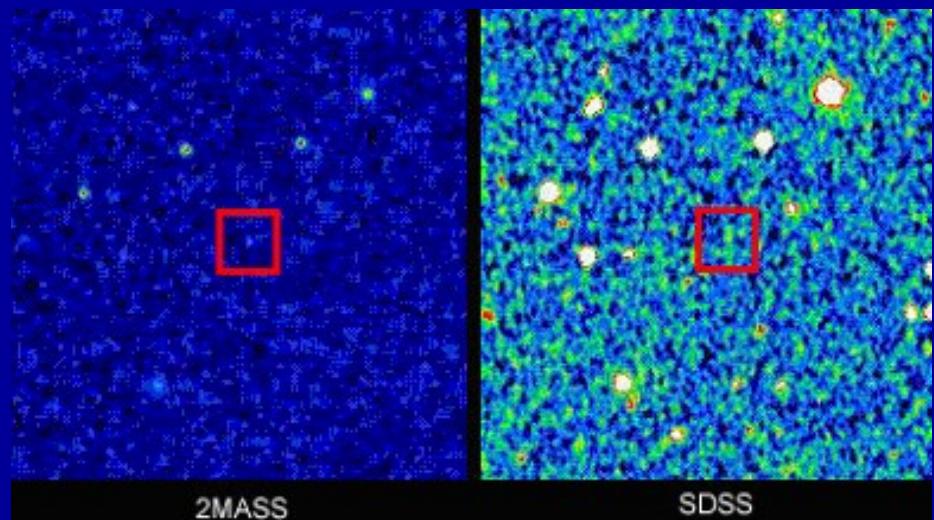
Scientific Motivation The search for brown dwarfs has been revolutionized by the latest deep sky surveys. A key attribute to discovering brown dwarfs is the federation of many surveys over different wavelengths. Such matching of catalogs is currently laborious and time consuming. This matching problem is generic to many areas of astrophysics.

Data Resources

- Sloan Digital Sky Survey (SDSS) Early Data Release (15 million objects)
- 2-Micron All Sky Survey (2MASS) 2nd Incremental Point Source Catalog (162 million objects)

What the VO Brings Today, doing the matching of these two large datasets is user-intensive and is replicated by many different users. Also, the correlation of these two datasets can take years of CPU time if not done correctly. The NVO brings two key aspects to

- **Filtering criteria:** z & J-only detections with $z-J > 2.75$
- *SDSS: 15M obj.*
- *2MASS: 160M obj.*
- *300000 objects in common.*



✓ *However, systematic searches using a VO methodology have not been performed so far.*

Democratization of Science

- **Digital Divide**
technological barrier, data access free, access to journals
- **International Council for Science CODATA**
Committee on Data for Science and Technology – UNO ICS
- **CASPAR**
Cultural, Artistic and Scientific knowledge for Preservation, Access and Retrieval
Digital curation centers
- **ADS and VO**
links to ivo://, metadata, ontologies – semantic web
- **Archive importance:**
5x IUE , 3x HST results from archives than PI articles
- **Effectiveness**
50% of published data appears in Journals, links to data automatic ?

Astroinformatics

- Analogy – Bioinformatics (Genome analysis with GRIDS, ATB)
- e-Science in Astronomy - using informatics (computer science)
- 4-th Paradigma of science (observation, experiment, modeling, knowledge)
- Data mining, Knowledge discovery - VO-NEURAL, DAME
- Clustering
- Classification
- Supervised learning (Neural Networks, SVM)
- Examples
 - Photometric RedShift
 - Searching for QSO
 - Automatic Light curves classification (GAIA, LSST)
- Very NEW – emerging discipline

Objections to VO

Data quality – garbage in - garbage out

How and whom to give credit ? (button)

embedded ivo:// data in ApJ

VO for dissemination only

technology for OPTICON, nextgen

Virtual science – VO technology

VO only for public data ! Proprietary ?

(data jealousy)

local archive - available data marked

The Astronomer's Data Manifesto

at 26 IAU GA Prague SPS3

- (a) All significant tables, images, and spectra published in journals should appear in astronomical data centres.
- (b) All data obtained with publicly-funded observatories should, after appropriate pro-prietary periods, be placed in the public domain.
- (c) In any new major astronomical construction project, the data processing, storage, migration, and management requirements should be built in at an early stage of the project plan, and costed along with other parts of the project.
- (d) Astronomers in all countries should have the same access to astronomical data and information.
- (e) Legacy astronomical data can be valuable, and high-priority legacy data should be preserved and stored in digital form in the data centres.
- (f) The IAU should work with other international organisations to achieve our common goals and learn from our colleagues in other fields. ”

Czech VO - CZVO

Search | Login



CZVO
Czech Virtual Observatory

Navigation

- ▶ About CZVO
- ▶ Observatories
- ▶ Projects
- ▶ Spectra Archive
- ▶ Data Resources
- ▶ Links
- ▶ Publications

Home

Links

There is some links:

- Virtual Observatory United Kingdom
- Astrogrid
- Australian Virtual Observatory/
- Chinese Virtual Observatory
- Canadian Virtual Observatory
- European Virtual Observatory
- German Astrophysical Virtual Observatory
- Hungarian Virtual Observatory
- Japanese Virtual Observatory
- Korean Virtual Observatory
- National Virtual Observatory, United States
- Observatoire Virtuel France
- Russian Virtual Observatory
- Spanish Virtual Observatory
- Italian Virtual Observatory
- Virtual Observatory India

Powered by [Drupal](#) - Design by [SKooDA](#)



CZVO Activities

VO-KOREL (web services)

parallel run of many jobs – more users
using VO Universal Worker Server (CEA)
job control, queuing, jobs results polling
will be integrated in VODesktop

1D spectra cutout server (HEROS)

SSA access to 1D spectra + cutout of regions (lines)
normalization, (rebinning , convolution on server)

Data mining – AstroNeural + Clustering

VIRTUAL OBSERVATORY

